
Characterization of discharge based plasmas in the spectral range of 20-50 nm

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CONTENT

- Motivation for 20 nm – 50 nm
- Experimental set-up
- Results on emission for different gases
- Conclusions

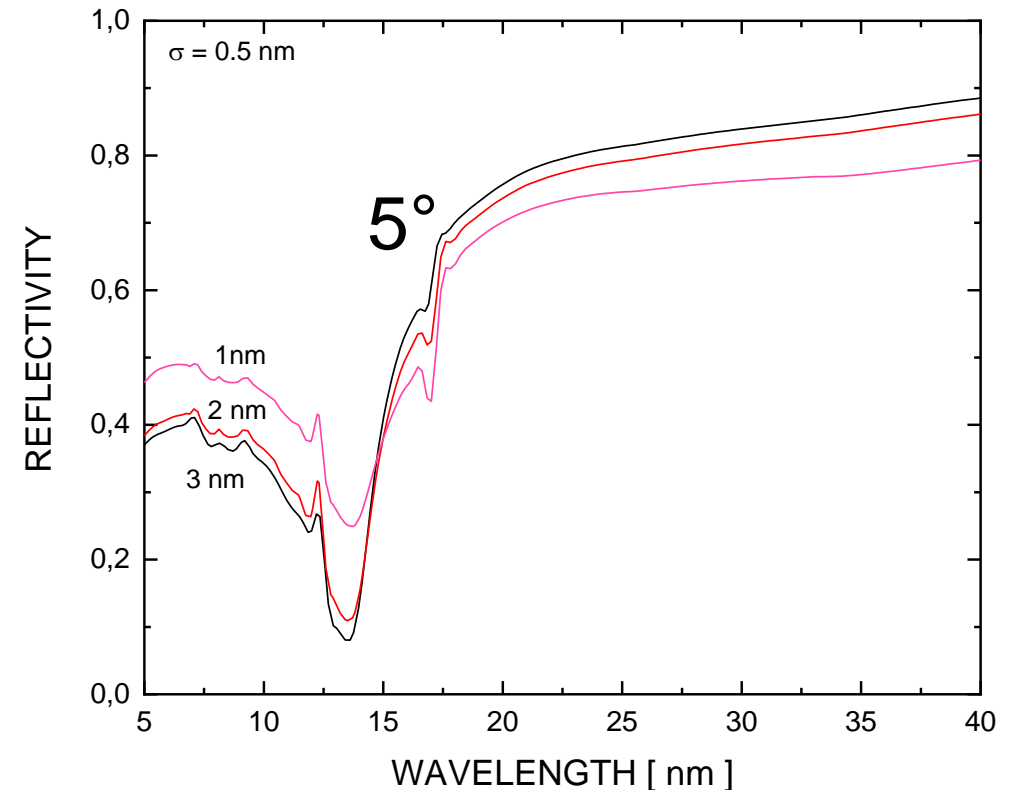
Why dealing with 20-50 nm ?

- higher relative Mie scattering compared to 193 nm, benefit for defect inspection using 47 nm¹
- increase of reflectivity with wavelength coming from lower wavelengths
- higher reflectance angle, better spatial resolution larger data set for reconstruction of parameters
 - characterization of layer systems
 - spectroscopic reflectometry for periodic structure characterization²

¹Ref.: Barnes et al., SPIE Advanced Lithography 2017, Metrology, Inspection and Process control for Microlithography XXXI, Vol. 10145, 2017

²Ref.: Bahrenberg et al., SPIE Advanced Lithography 2019, Metrology, Inspection and Process control for Microlithography XXXIII, Vol. 10959, 2019

Determination of Al₂O₃ thickness on silicon substrate



Ref.: CXRO – The Center for X-Ray Optics
(<http://www.cxro.lbl.gov/>)

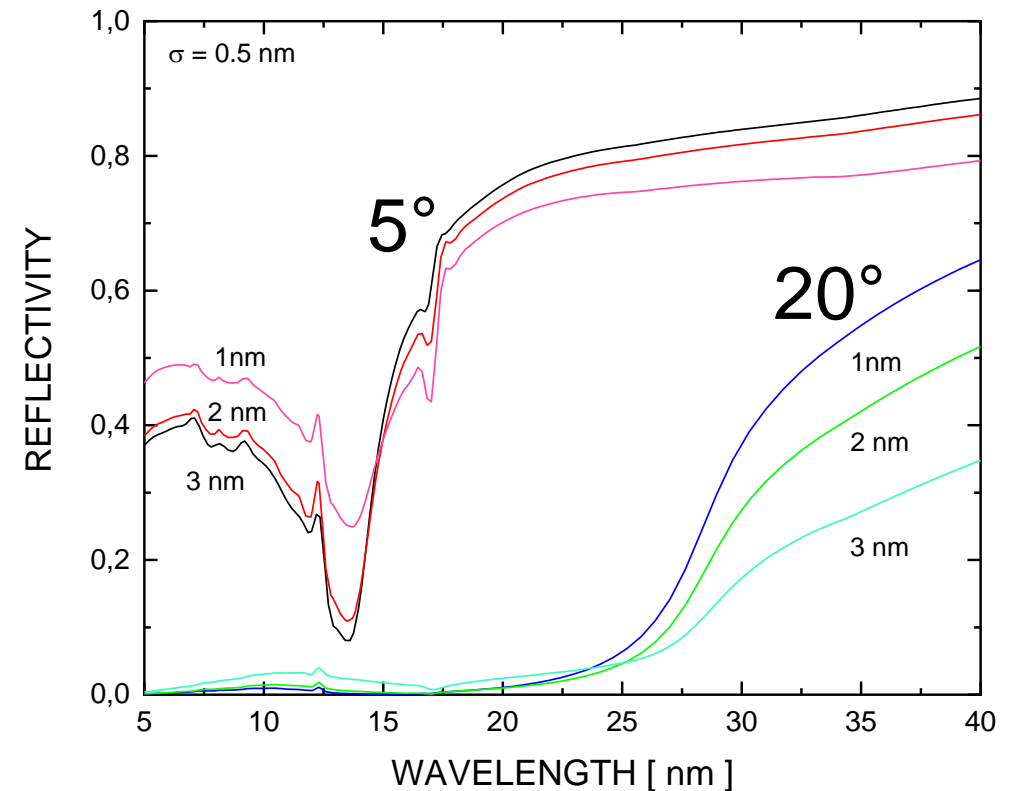
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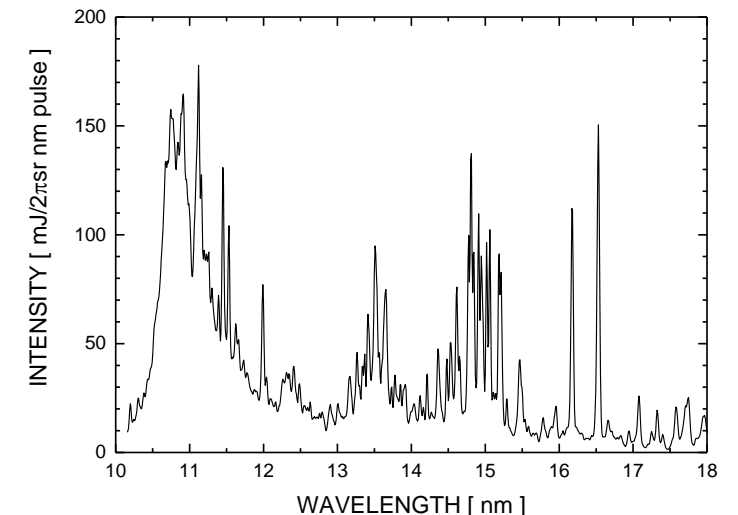
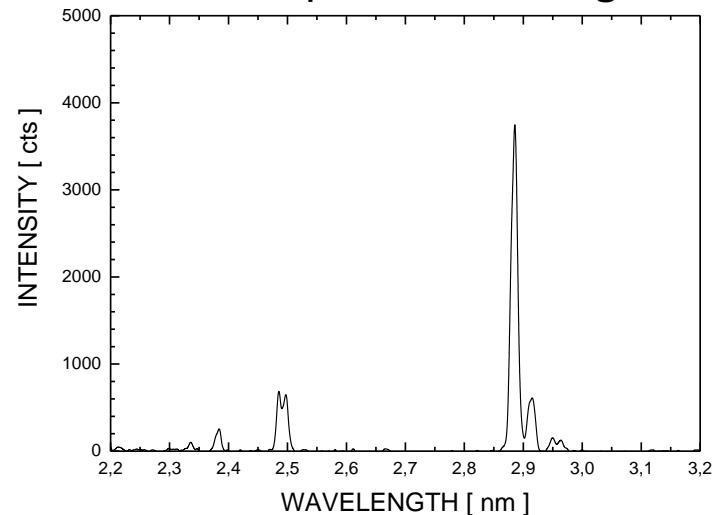
Opportunities and challenges with discharge plasmas

- Starting point for development:
 - pinch plasma source optimized for 2-20 nm
 - electrical input energy ~ 2-10 J
 - average input power up to 10 kW
- Optimized emission for 20-50 nm is expected at lower plasma temperature/lower input energy
 - lower electrode erosion
 - longer maintenance interval
 - lower Cost of Ownership
- Planck radiator:
lower temperature means
broader wavelength interval
- Technical challenge to produce
colder plasma with high density

Fraunhofer ILT
XUV source



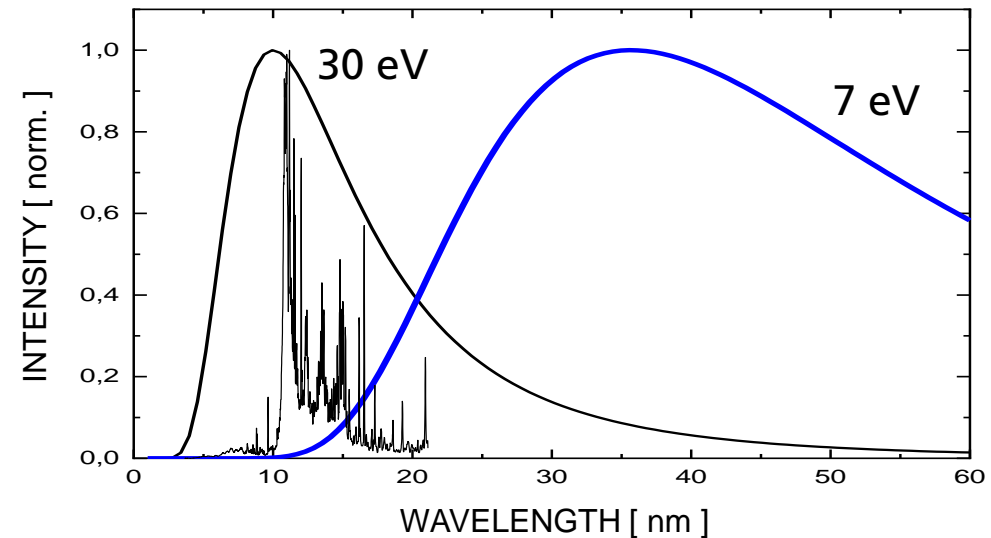
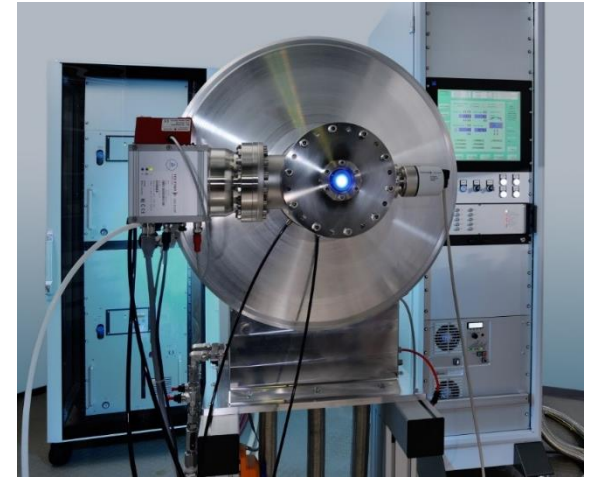
Emission spectra of nitrogen and xenon



Opportunities and challenges with discharge plasmas

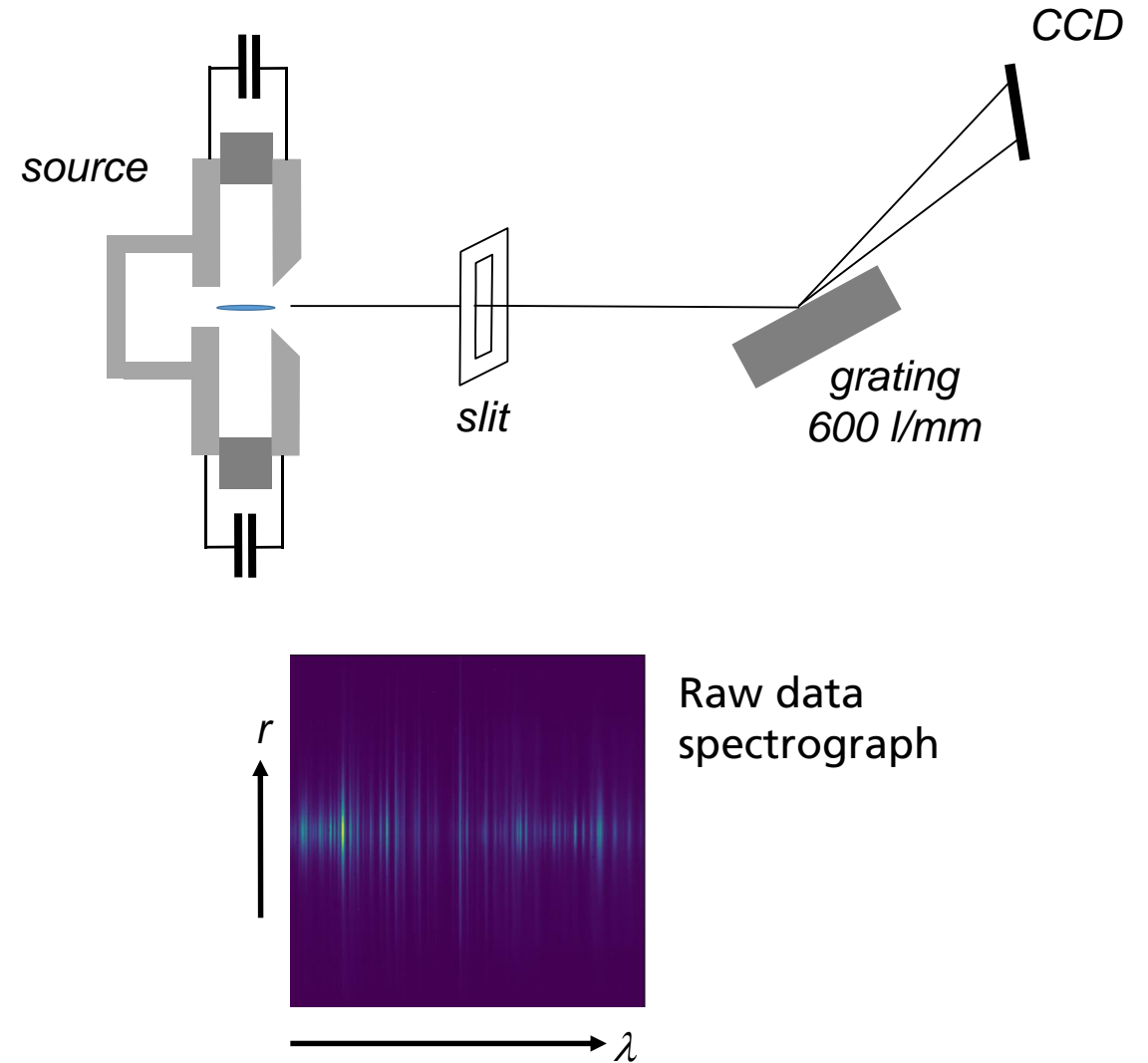
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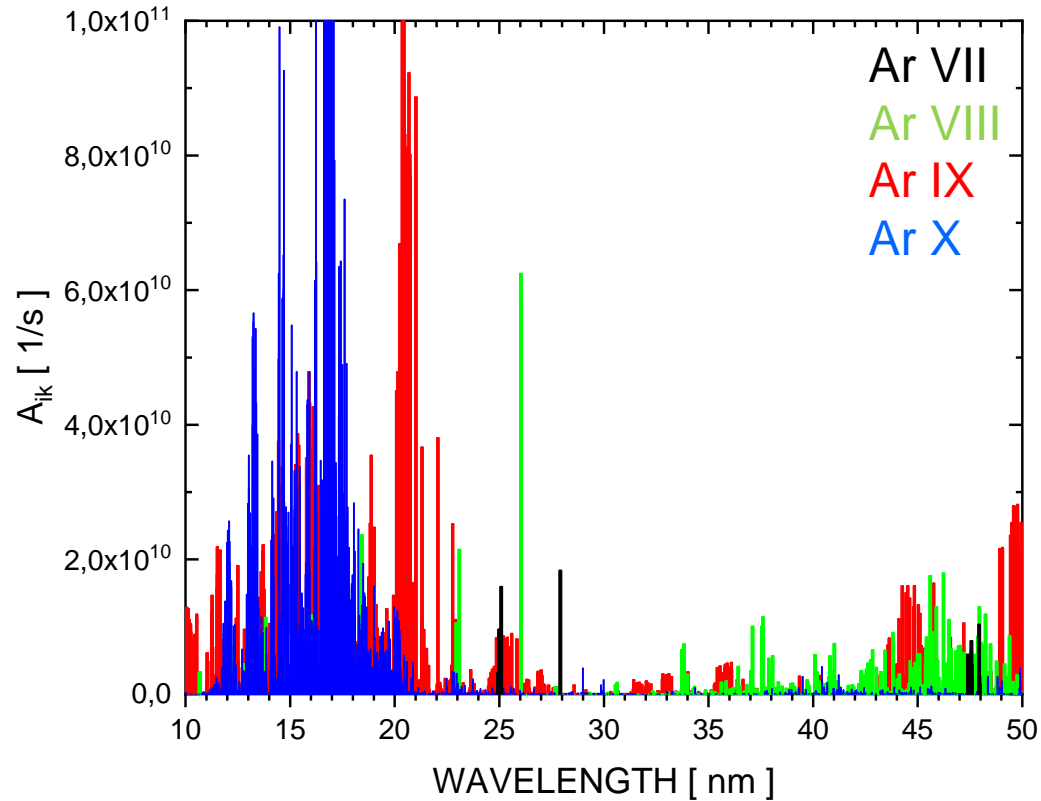
Experimental setup for 20 - 50 nm characterisation

- Hollow Cathode Triggered pinch plasma:
 - repetition rate 20 Hz - 30 Hz
 - input energy 1,3 J - 2,3 J
 - operation with Ne, Ar, Kr, Xe
- emission spectra recorded with grazing incidence spectrograph (600 lines/mm)
- spectrally resolved end-on emission profile using slit perpendicular to entrance slit of spectrograph (Abel inversion)
- absolute emission estimated with photo diodes + spectral filters



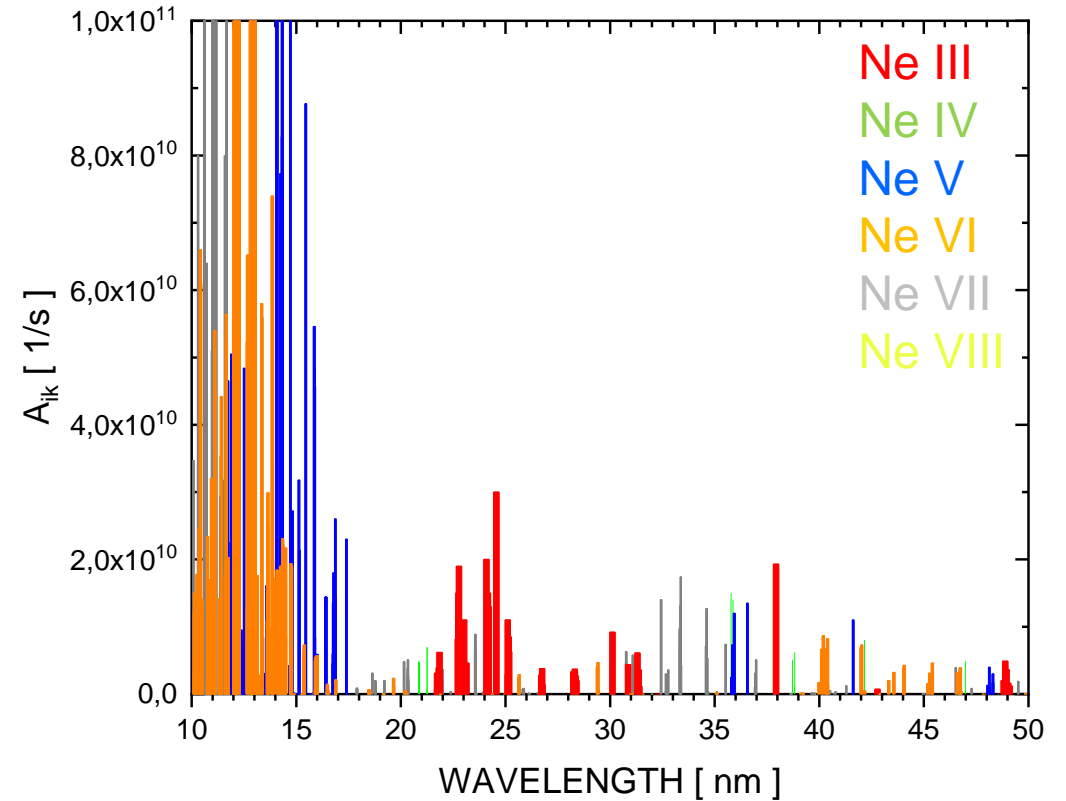
Atomic data

Argon



Ref.: CHIANTI - An Atomic Database for Spectroscopic Diagnostics of Astrophysical Plasmas (<https://chiantidatabase.org/sswidl/dbase/ar/>)

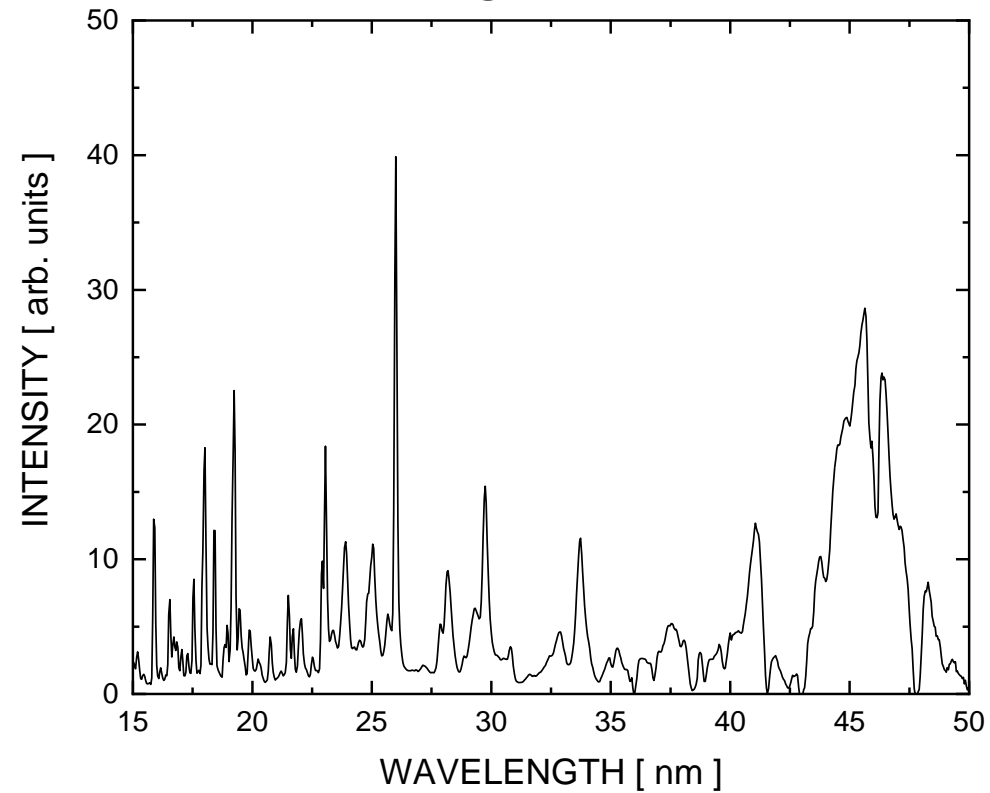
Neon



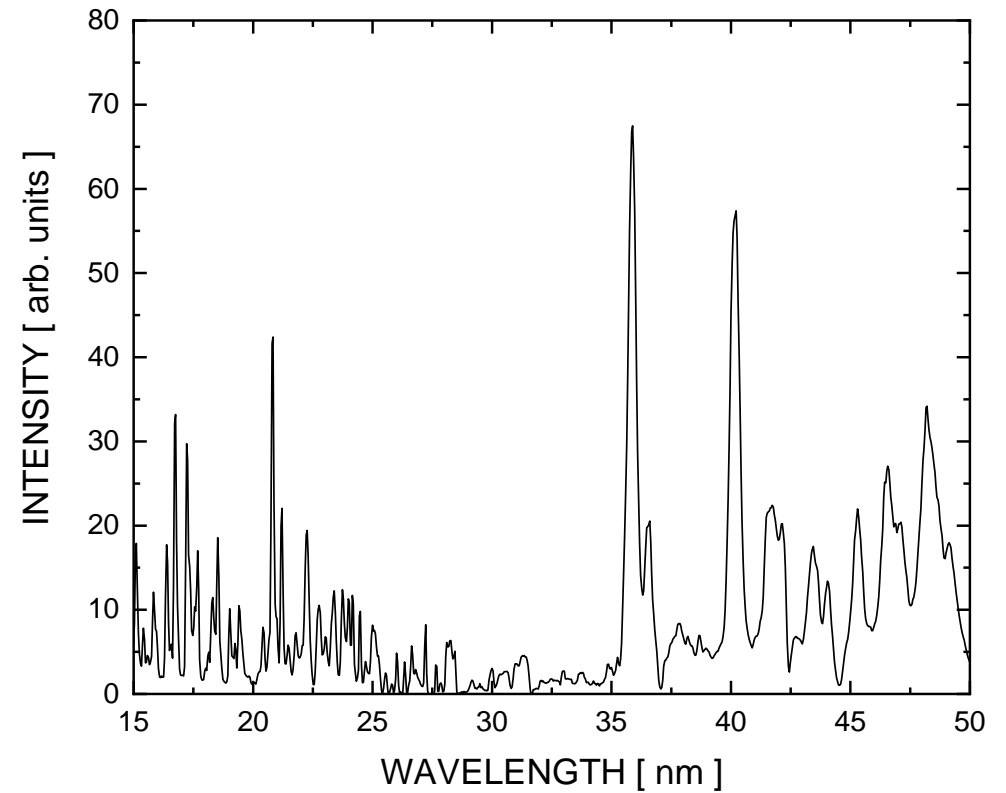
Ref.: National Institute of Standards NIST – Atomic Spectra Database (<https://nist.gov/pml/atomic-spectra-database>)

Emission spectra (1)

Argon

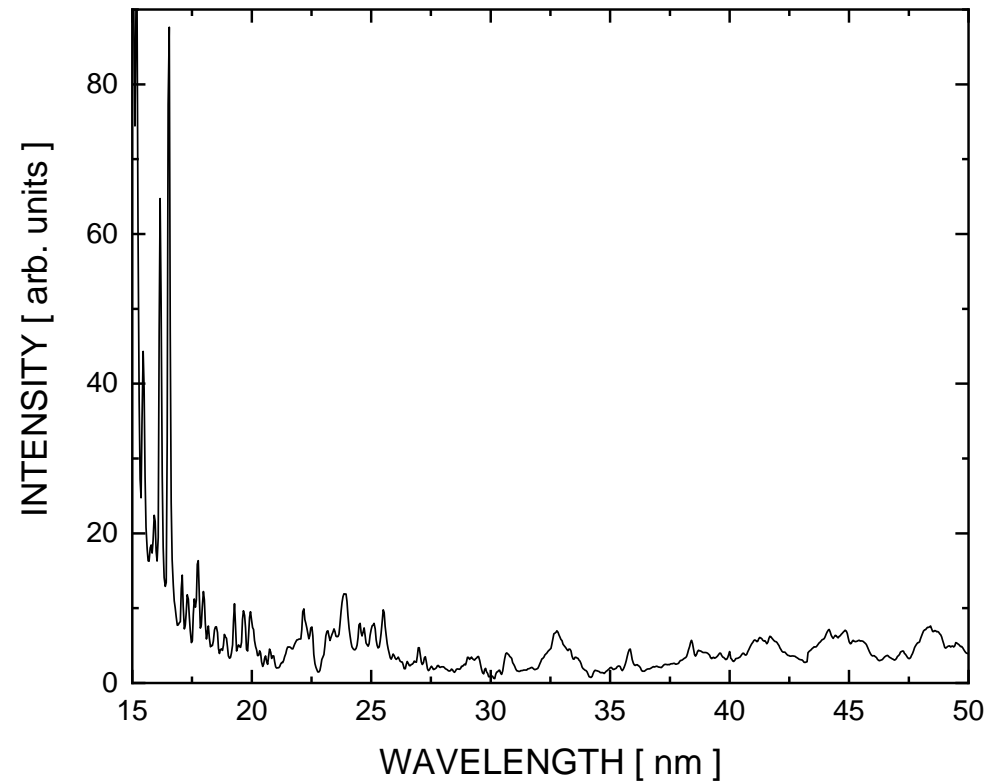


Neon

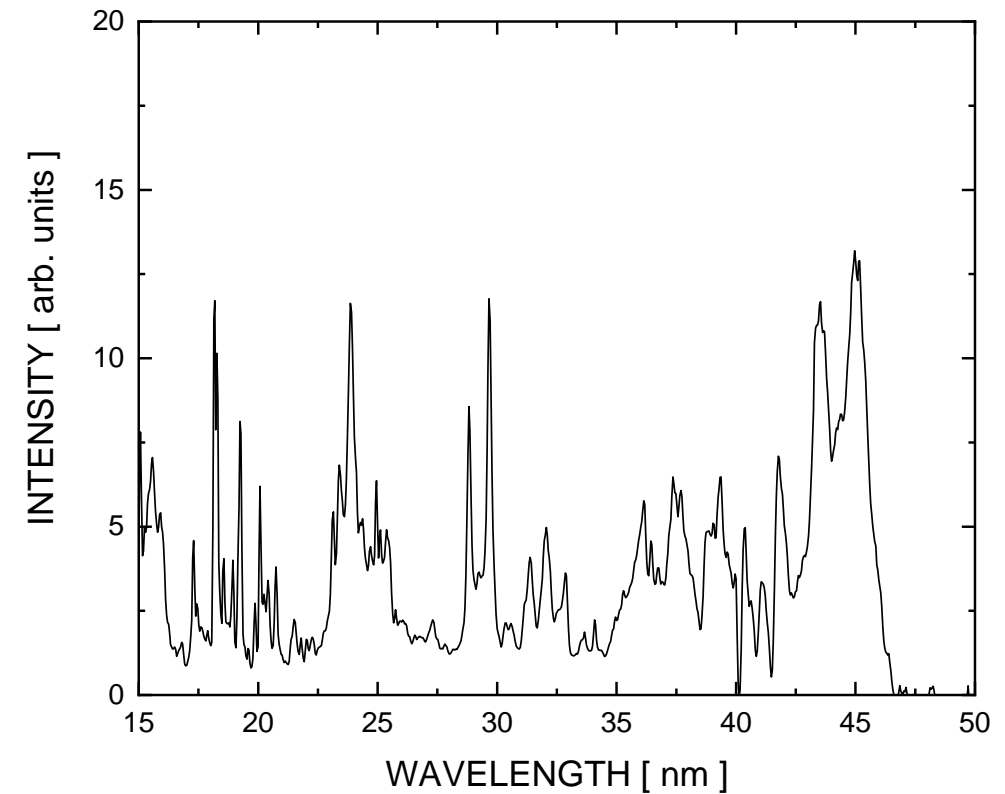


Emission spectra (2)

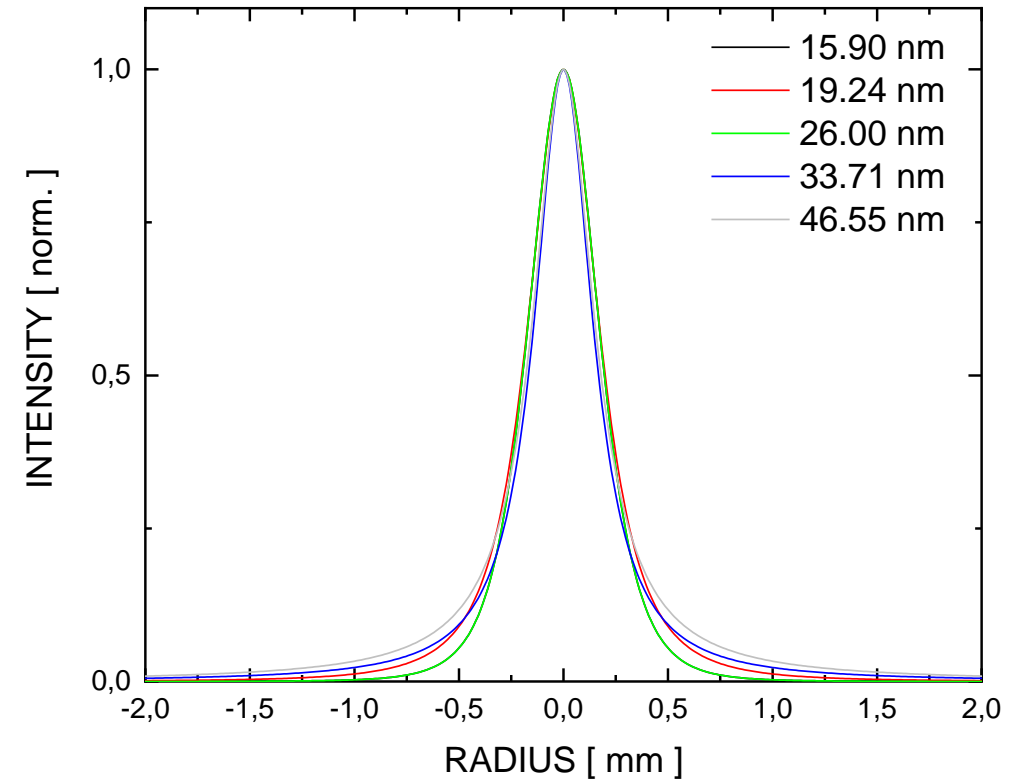
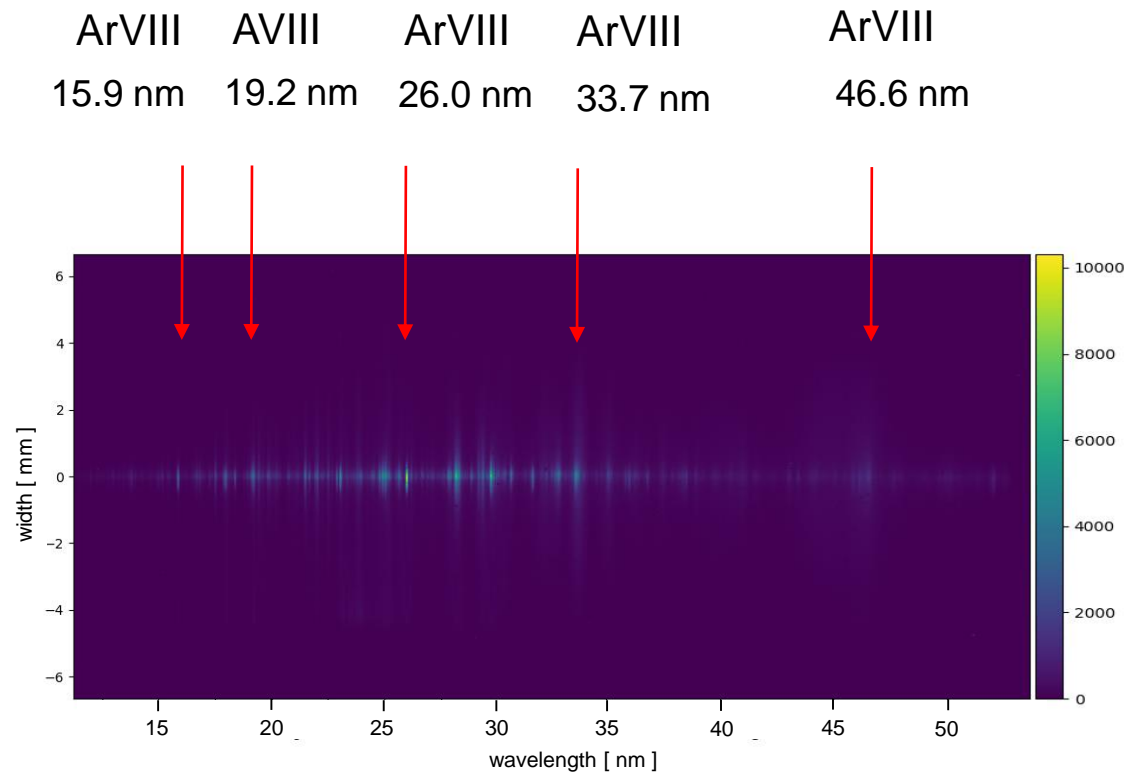
Xenon



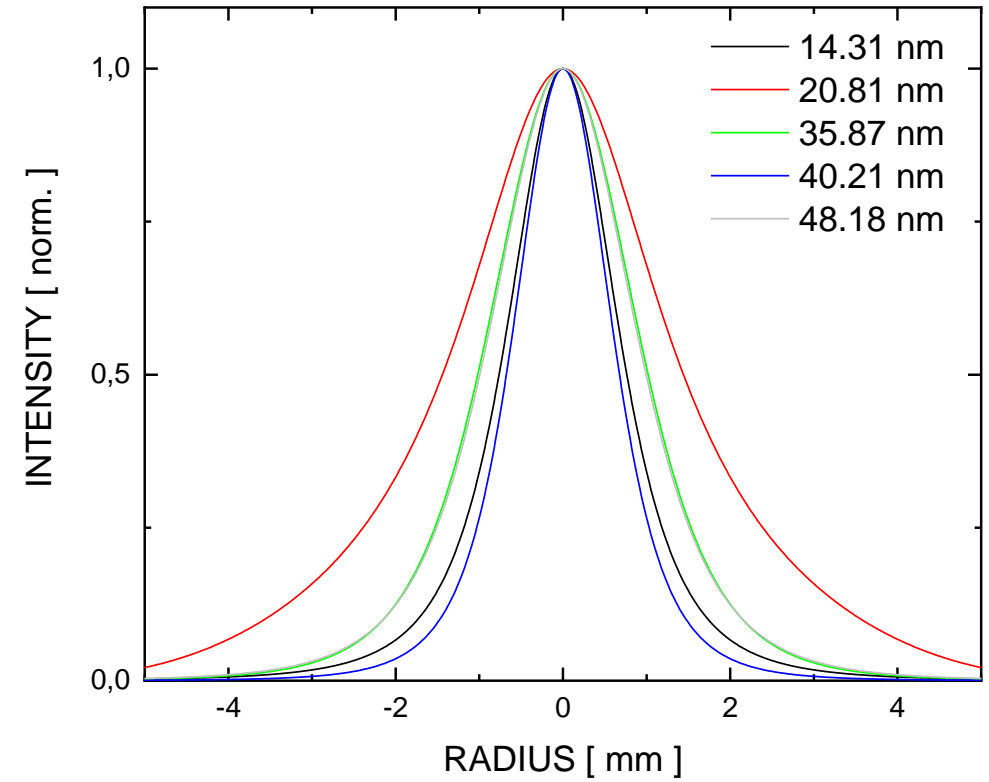
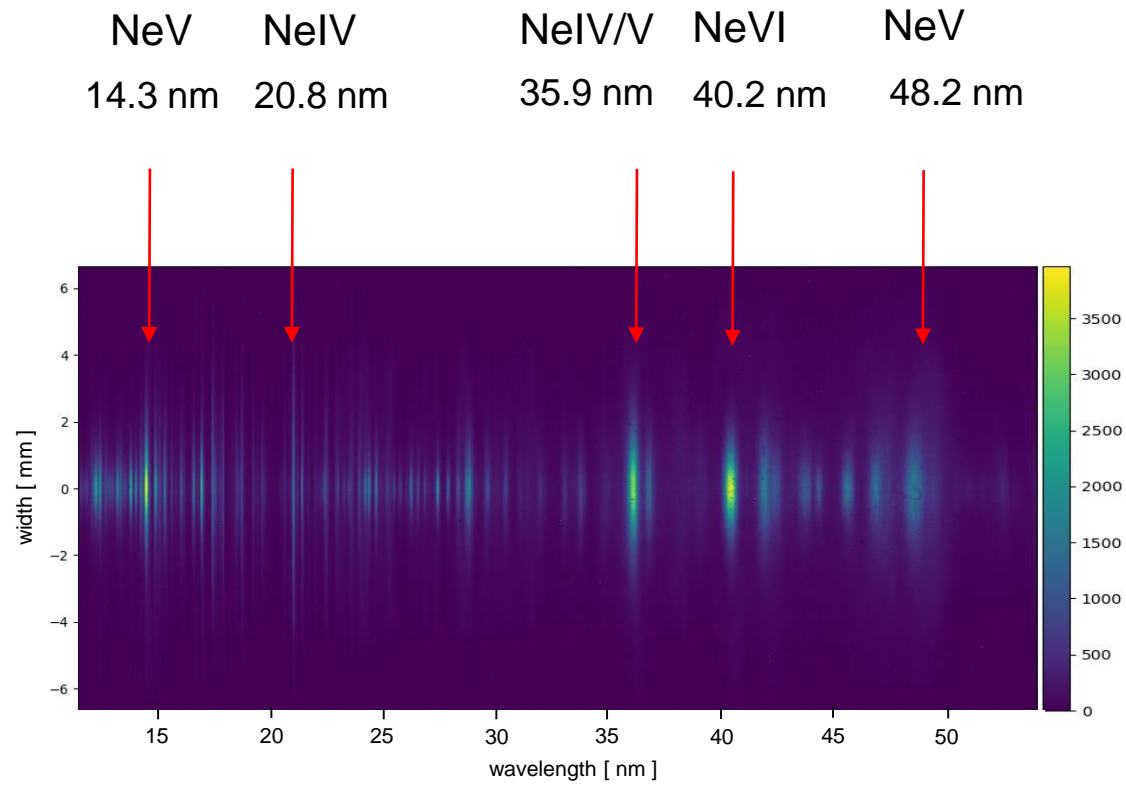
Krypton



Spatial profiles: Argon

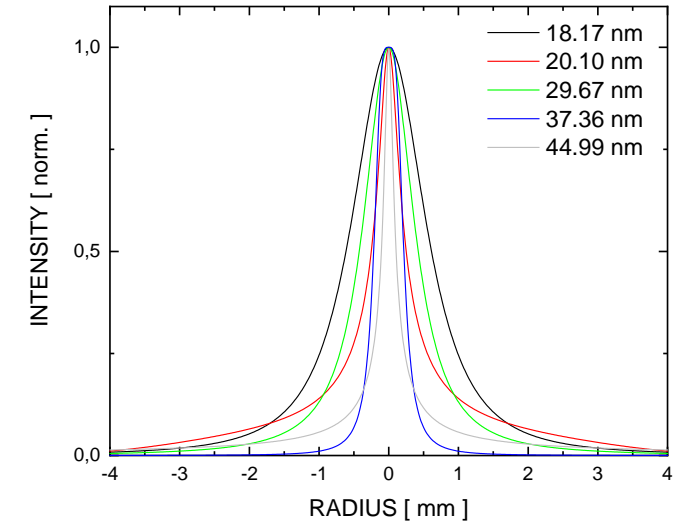
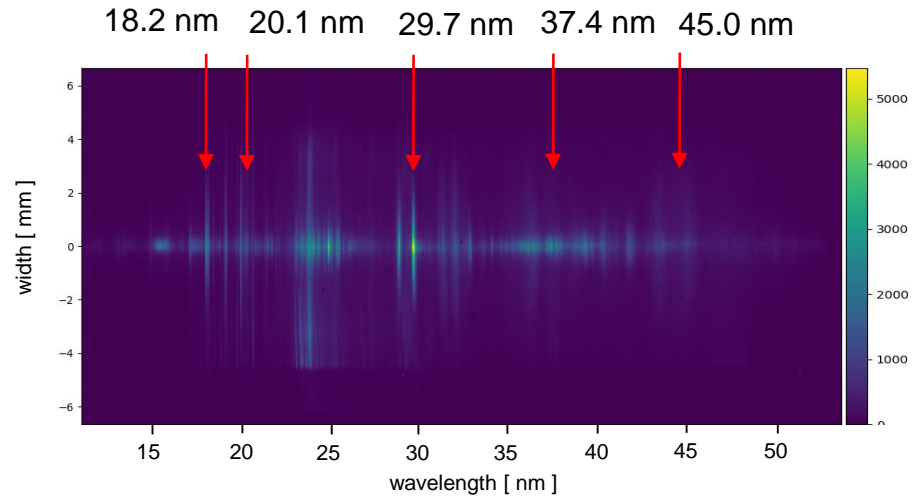


Spatial profiles: Neon

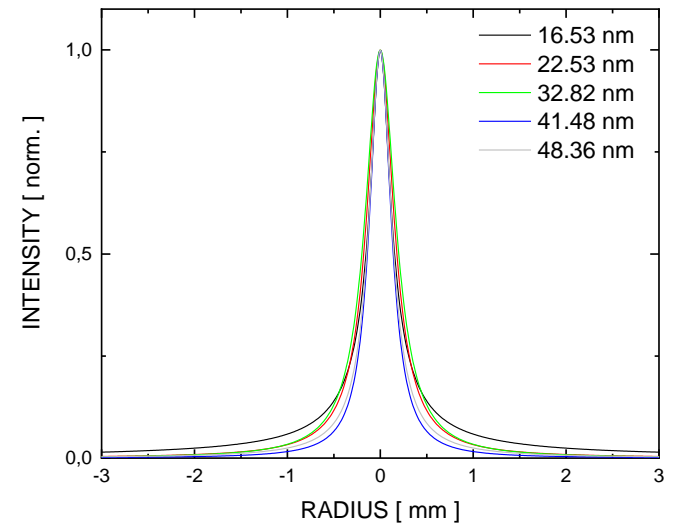
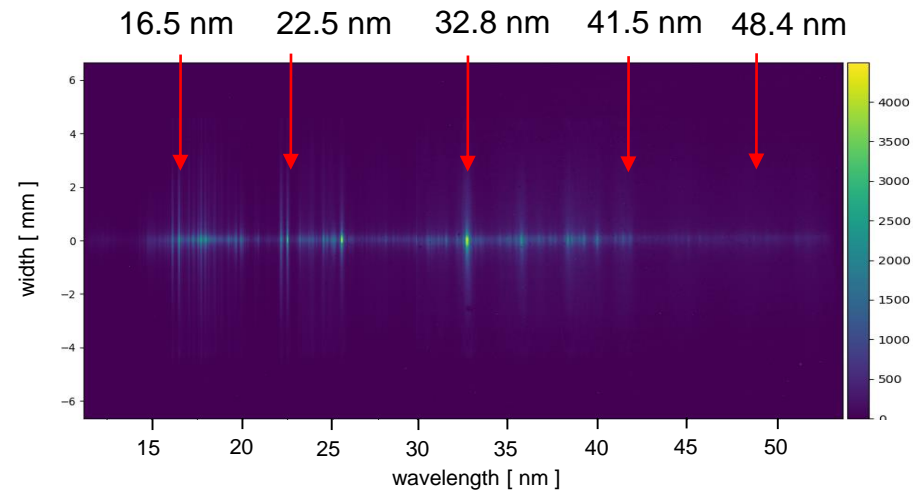


Spatial profiles: Krypton, Xenon

Krypton

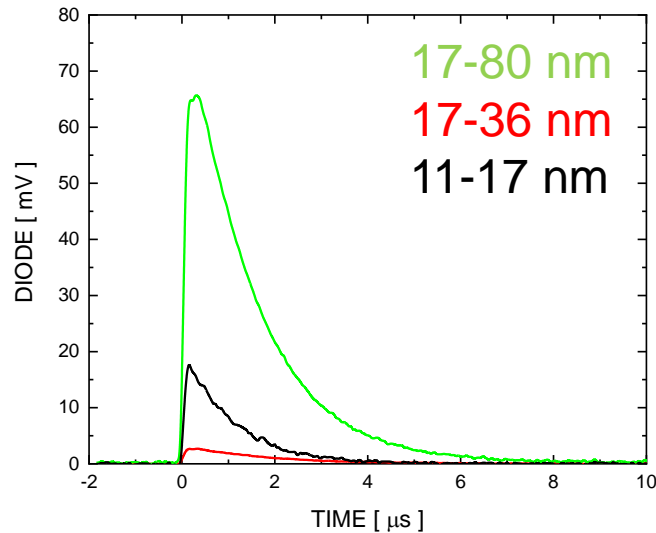
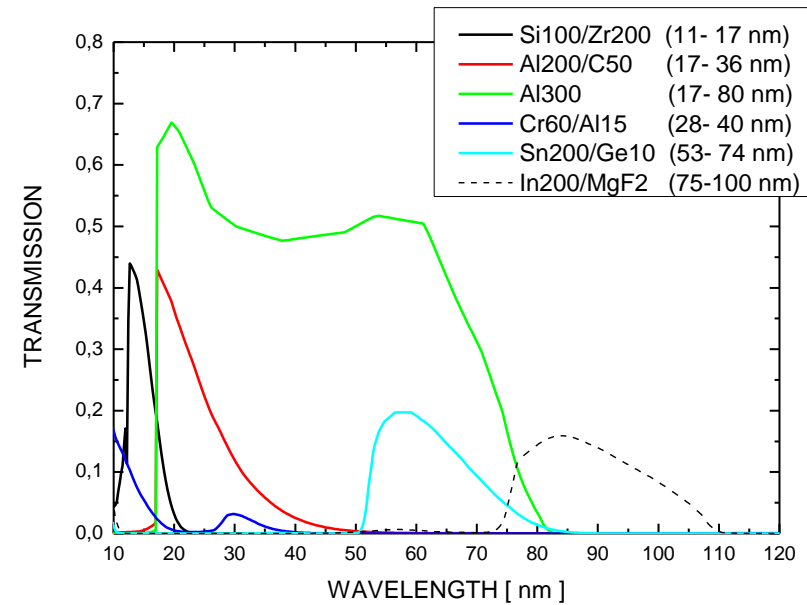


Xenon



Estimation of absolute emission

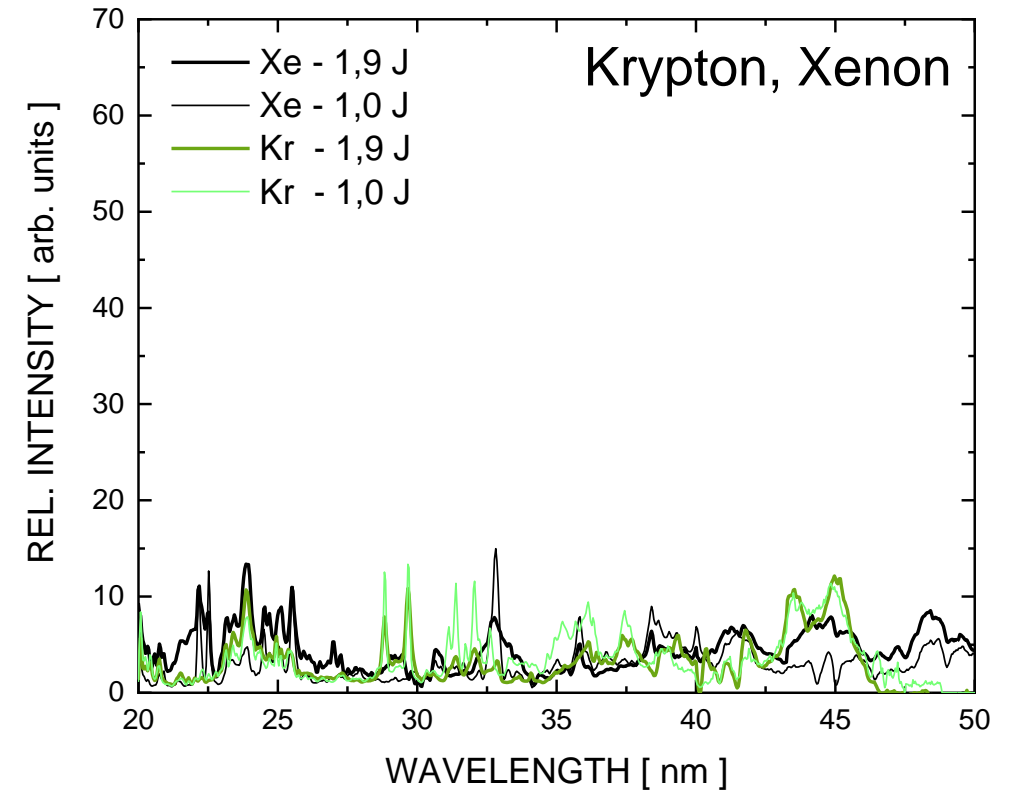
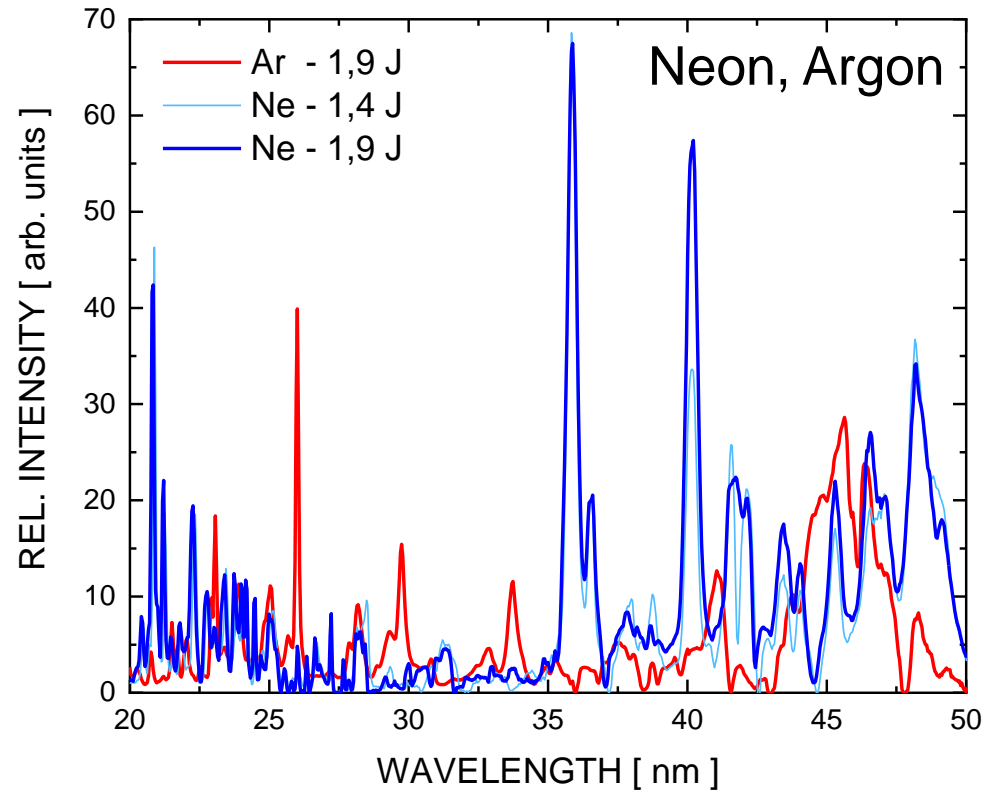
- nine AXUV 100 diodes
- sensitivity from suppliers data sheet ($\sim 0,2 \text{ A/W}$)
- different coatings for discrimination of spectral ranges
- correction for self-absorption in beamline



typical diode signals



Overview emission spectra



- emission spectra normalized to input energy (same scale for all elements !)
- expected total energy for Neon in wavelength interval 20-50 nm : $E_{20-50} \sim 100 \text{ mJ}/2\pi\text{sr}$
(rough estimation !!)
 $CE \sim 4 \%/2\pi\text{sr}$

Summary and conclusion

- Investigation on 20 nm - 50 nm emission of different gases in a ~2 J discharge plasma (Neon, Argon, Krypton, Xenon)
- Observed transitions of NeIV - NeVI and ArVIII
- “Large” range of diameter for different gases and for different ionization levels (up to 4 mm for NeIV)
- Highest conversion efficiency for Neon $\sim 100 \text{ mJ}/2\pi\text{sr}$ ($\sim 4 \text{ \%}/2\pi\text{sr}$)

see also poster S68:

Vieker et al., The Effect of Gas Admixture on the Operation of a Discharge based EUV-Source

THANK YOU !!