



# Drop impact phenomena in EUV lithography

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EUV source workshop 05-11-2019



Netherlands Organisation  
for Scientific Research

UNIVERSITY OF TWENTE.

TU/e

ASML



# Drop dynamics challenges in the EUV source

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- ❖ Optimal mass distribution for high CE

- ❖ Minimum debris

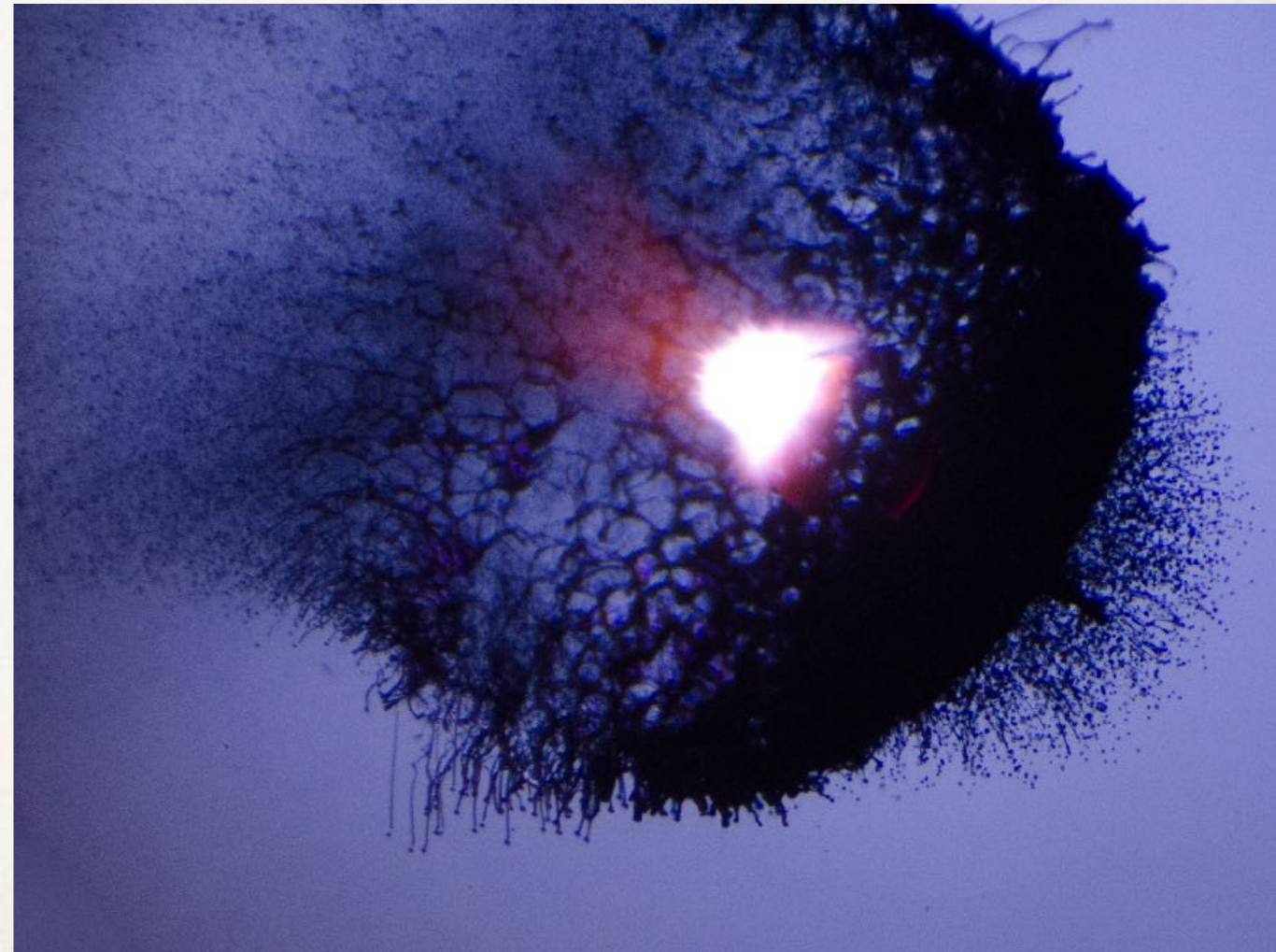
- ➔ How do drops deform?

- ➔ How do drops splash?

**Parameter space:**

prepulse energy, focus, beam profile, drop size, main pulse timing, wall temperature,..

**Required:** understanding basic fluid dynamics



# Drop impact in the EUV source

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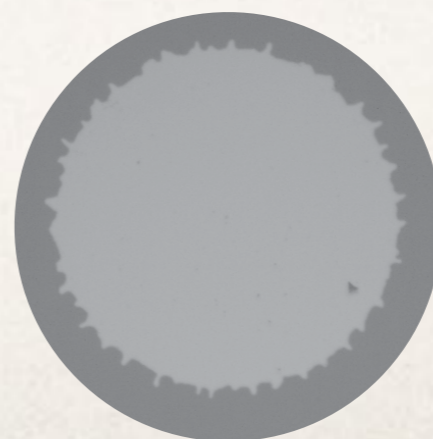
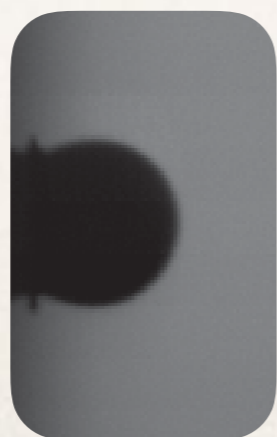
side view



front view

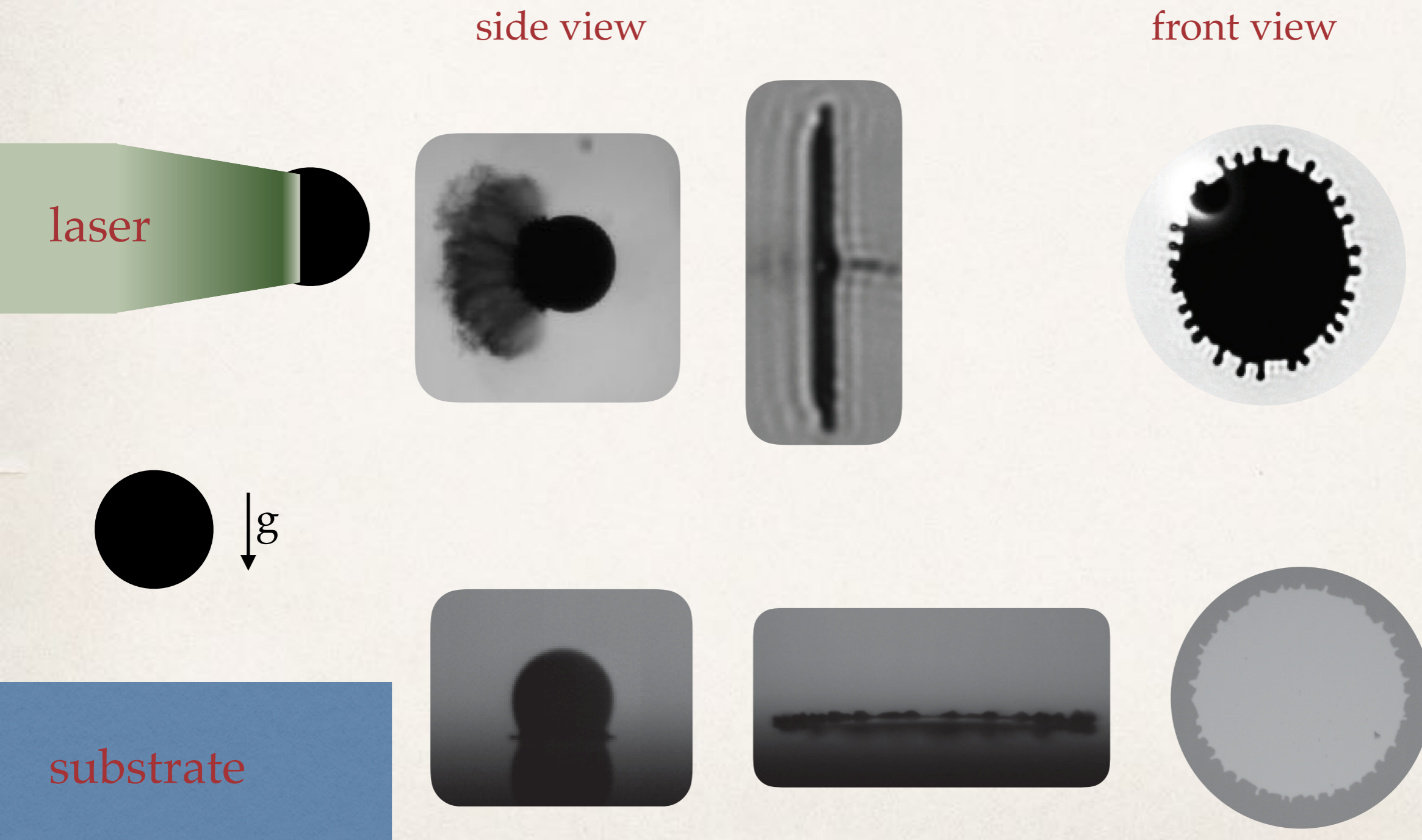


*What's the difference?*



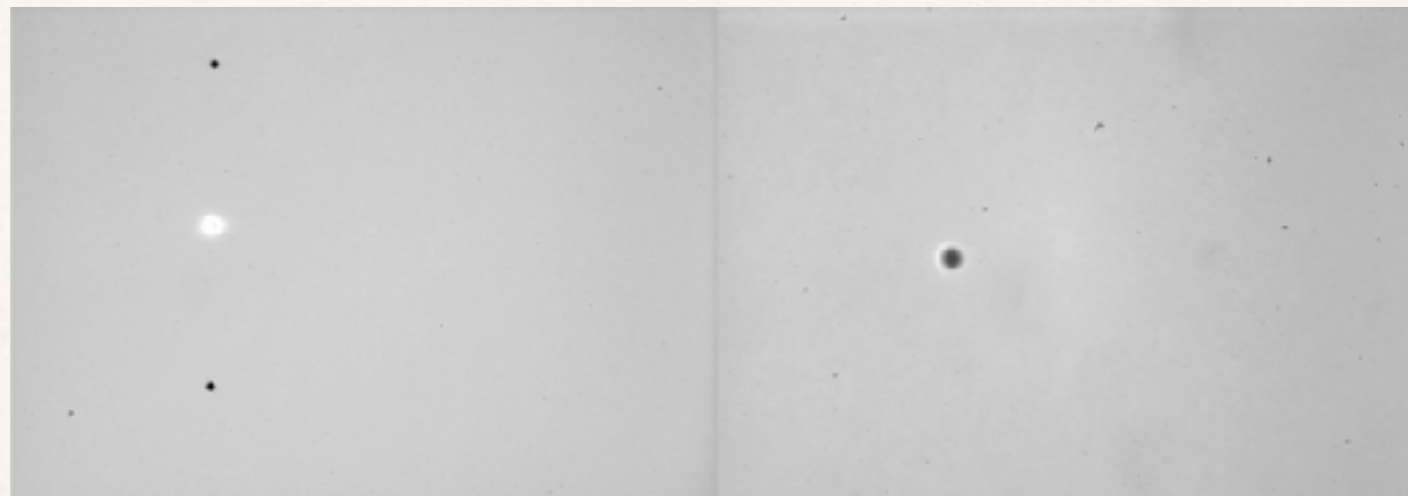
# Drop impact in the EUV source

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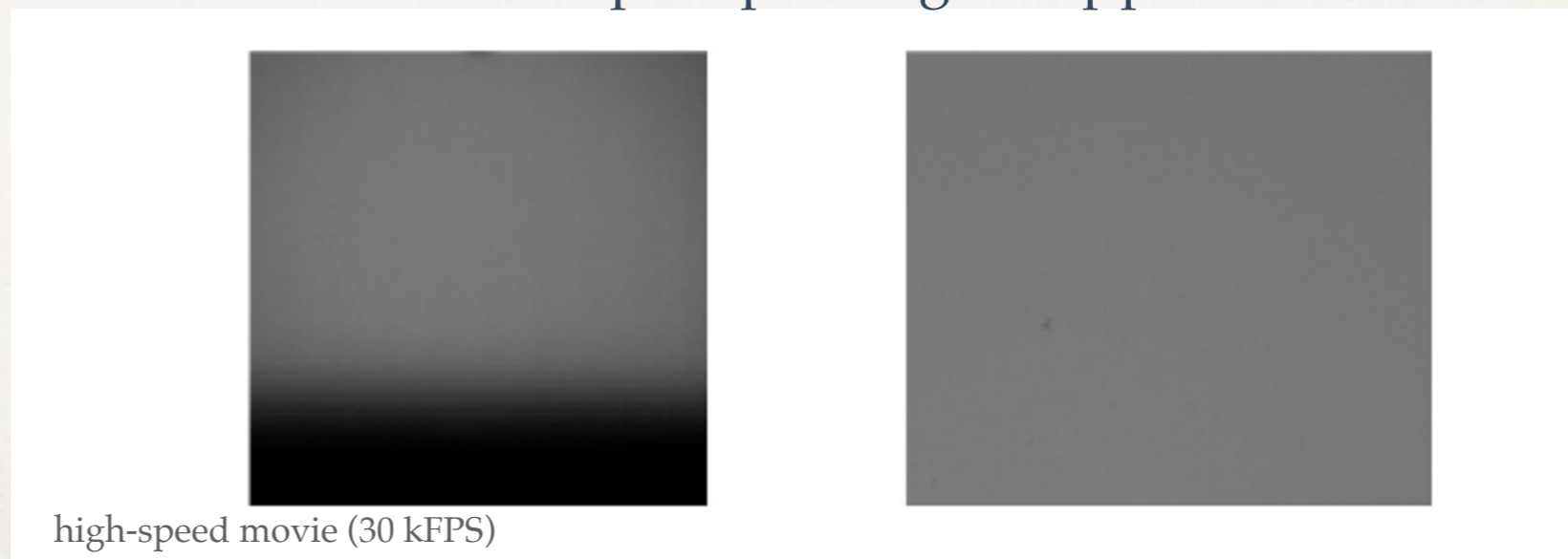
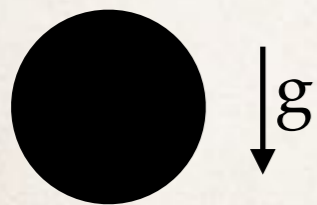
# Drop impact in the EUV source

1. A 25- $\mu\text{m}$  tin drop impacted by a 10 ns Nd:YAG laser



stroboscopic movie D. Kurilovich & O. Versolato (ARCNL)

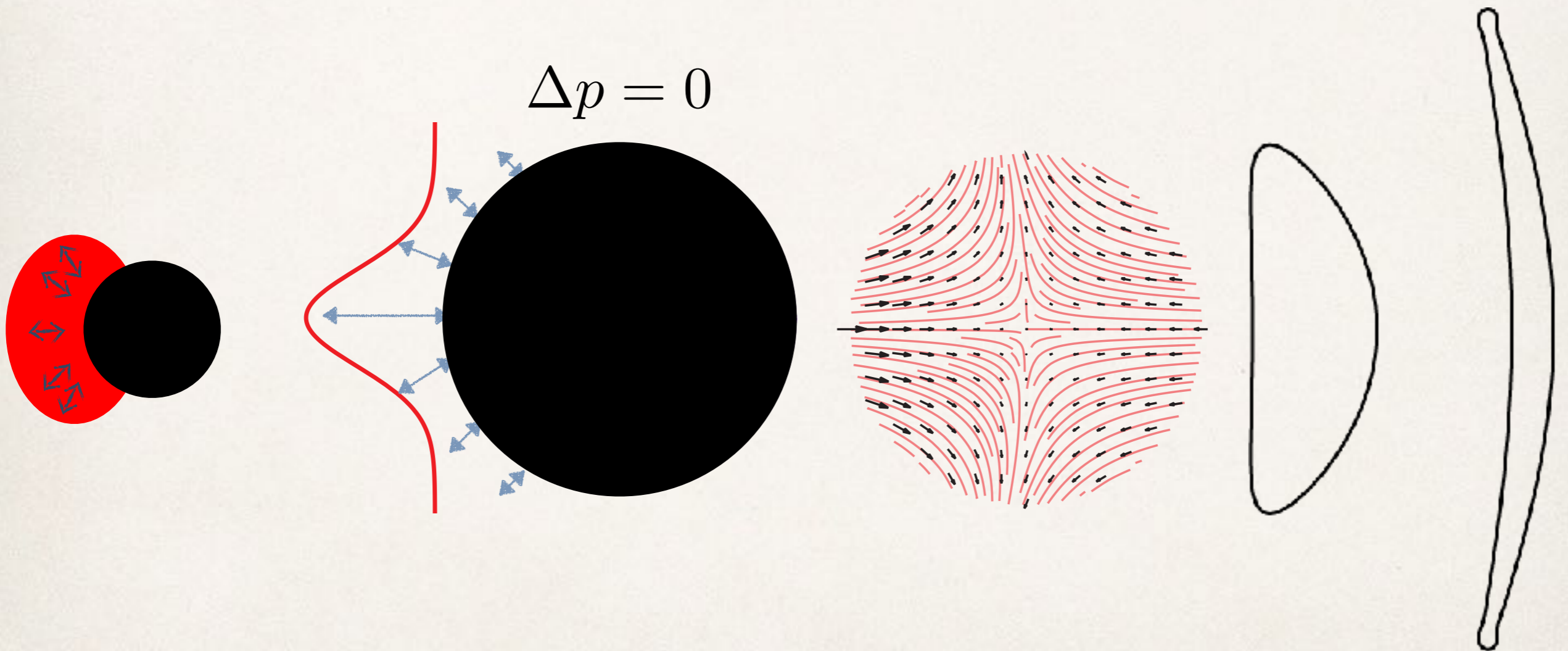
2. A 1.6-mm tin drop impacting a sapphire substrate



high-speed movie (30 kFPS)

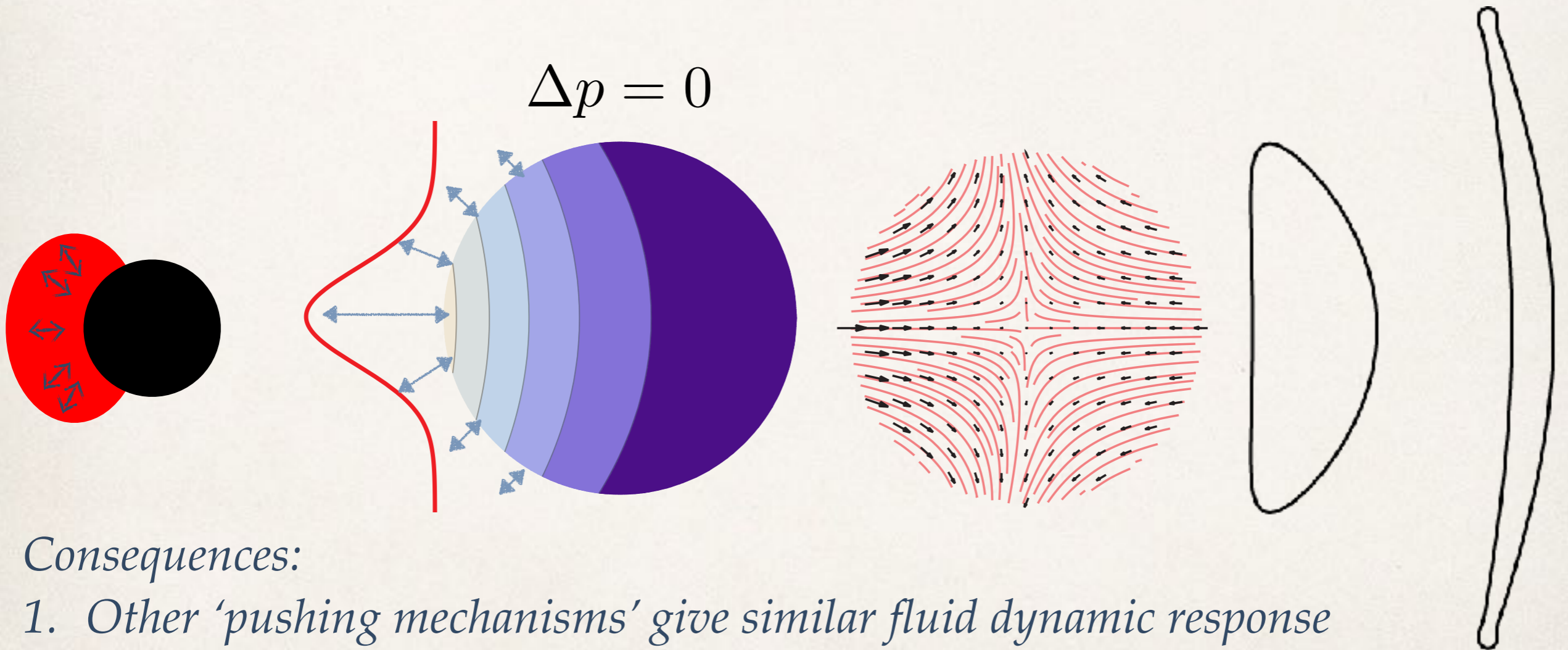
# Laser impact on a liquid drop: Pushing on a liquid sphere

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# Laser impact on a liquid drop: Pushing on a liquid sphere

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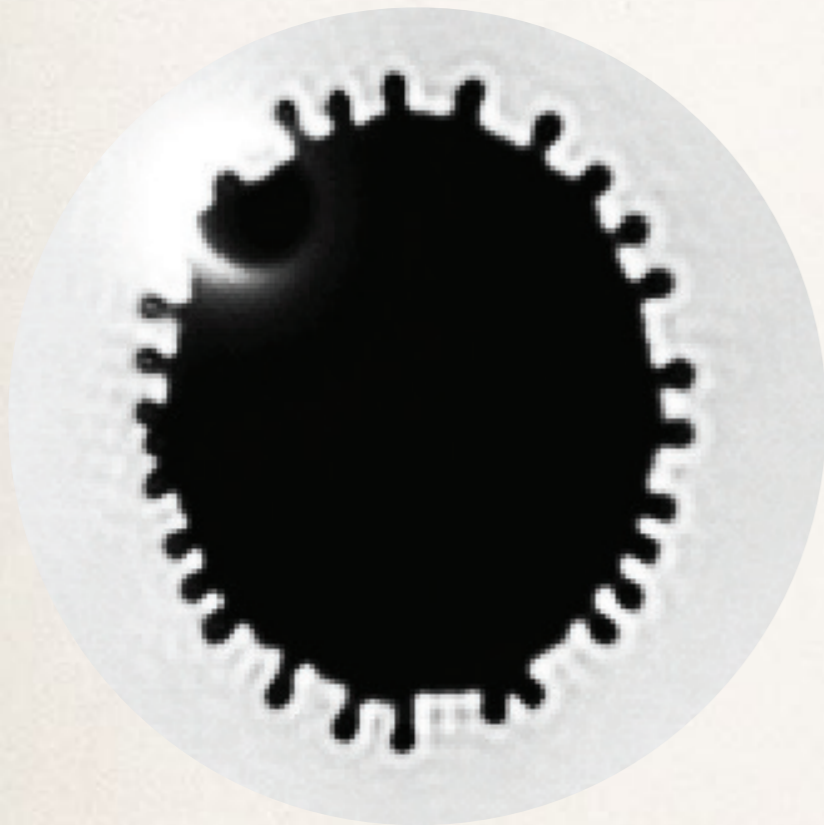
*Consequences:*

- 1. Other 'pushing mechanisms' give similar fluid dynamic response*
- 2. Pressure profile predicts direction of motion, shape & breakup*

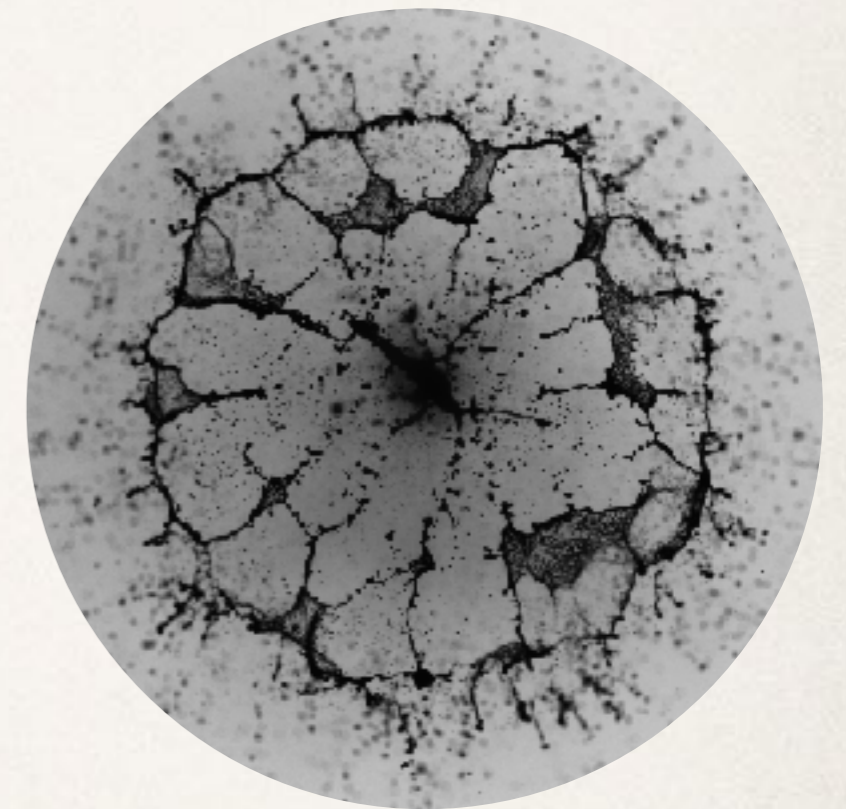
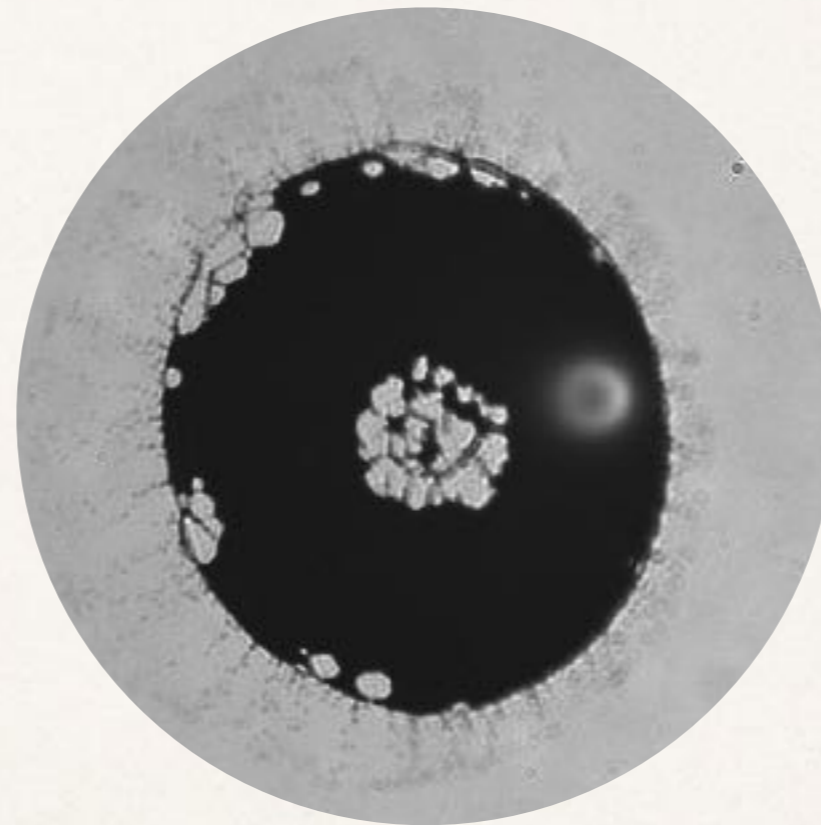
# Drop breakup scenarios

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rim breakup



hole nucleation



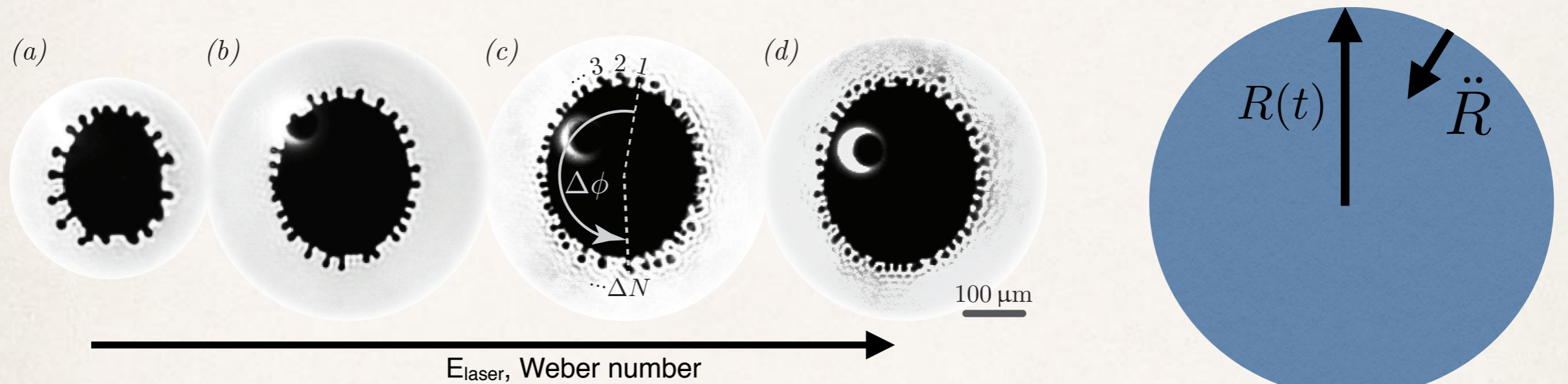
key parameter governing drop deformation & fragmentation

$$\text{Weber number} = \frac{\text{kinetic energy}}{\text{surface energy}} = \frac{\rho R_0 U^2}{\gamma}$$



# Rim breakup

## Rayleigh-Taylor & Rayleigh-Plateau instabilities



Predictions:

$$\omega \sim \left( \frac{\rho \ddot{R}^3}{\gamma} \right)^{1/4}$$

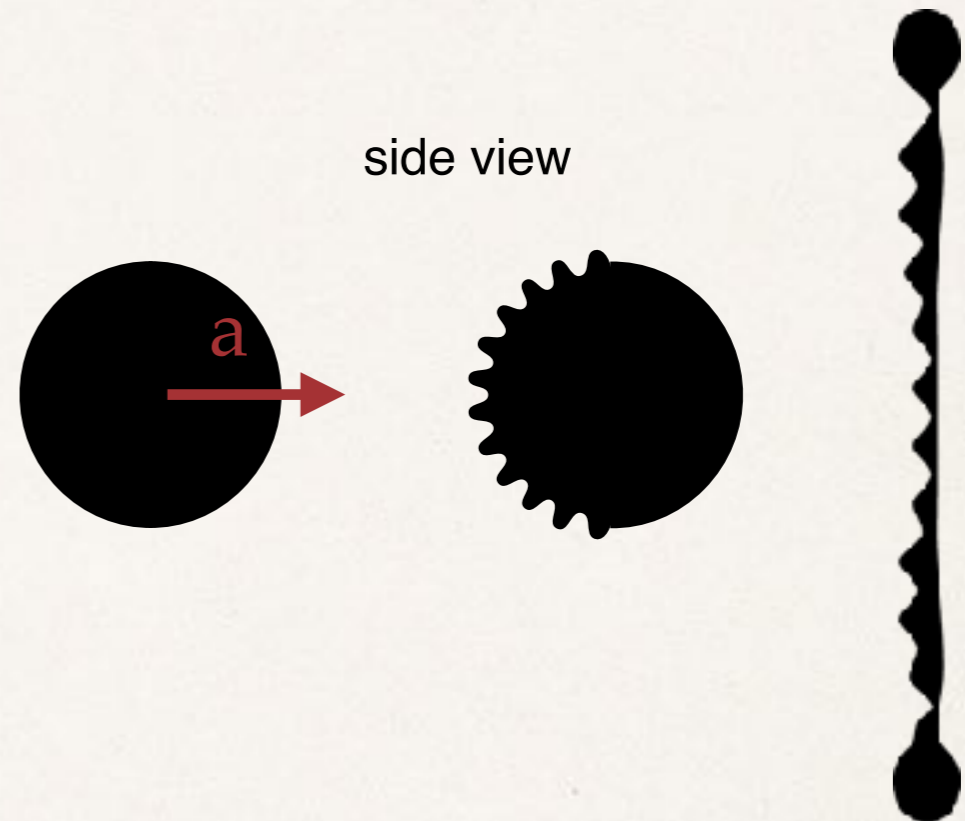
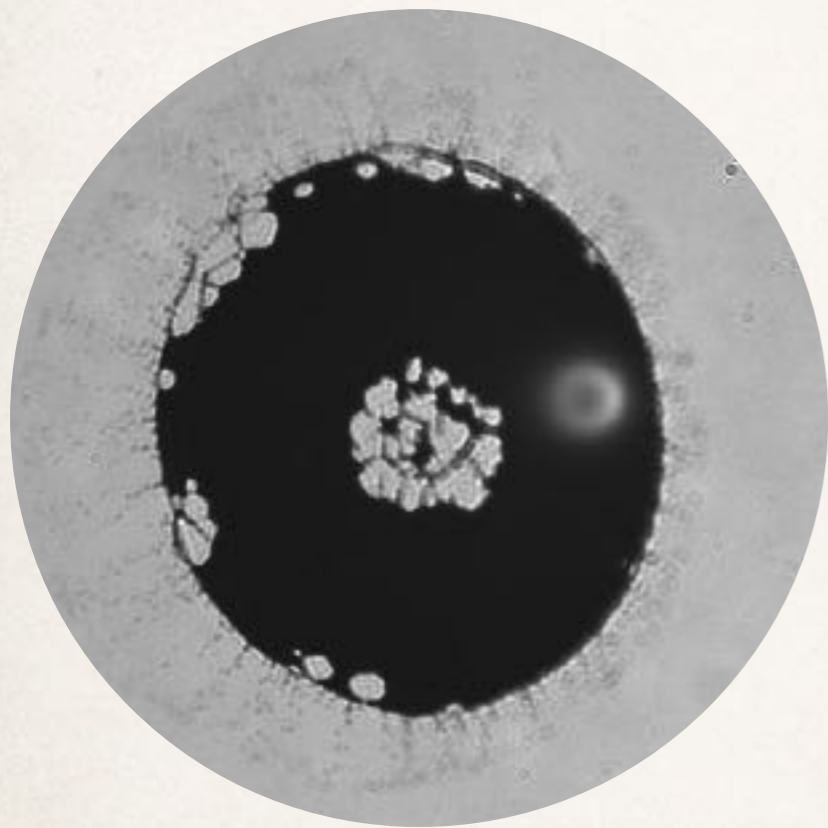
$$\lambda \sim \left( \frac{\gamma}{\rho \ddot{R}} \right)^{1/2}$$



# Hole nucleation

RT instability caused by impulsive acceleration

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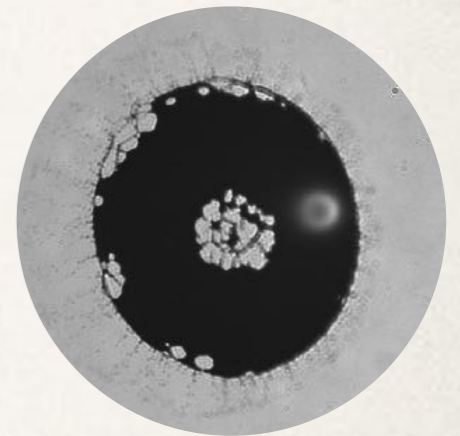


# Prediction of breakup time



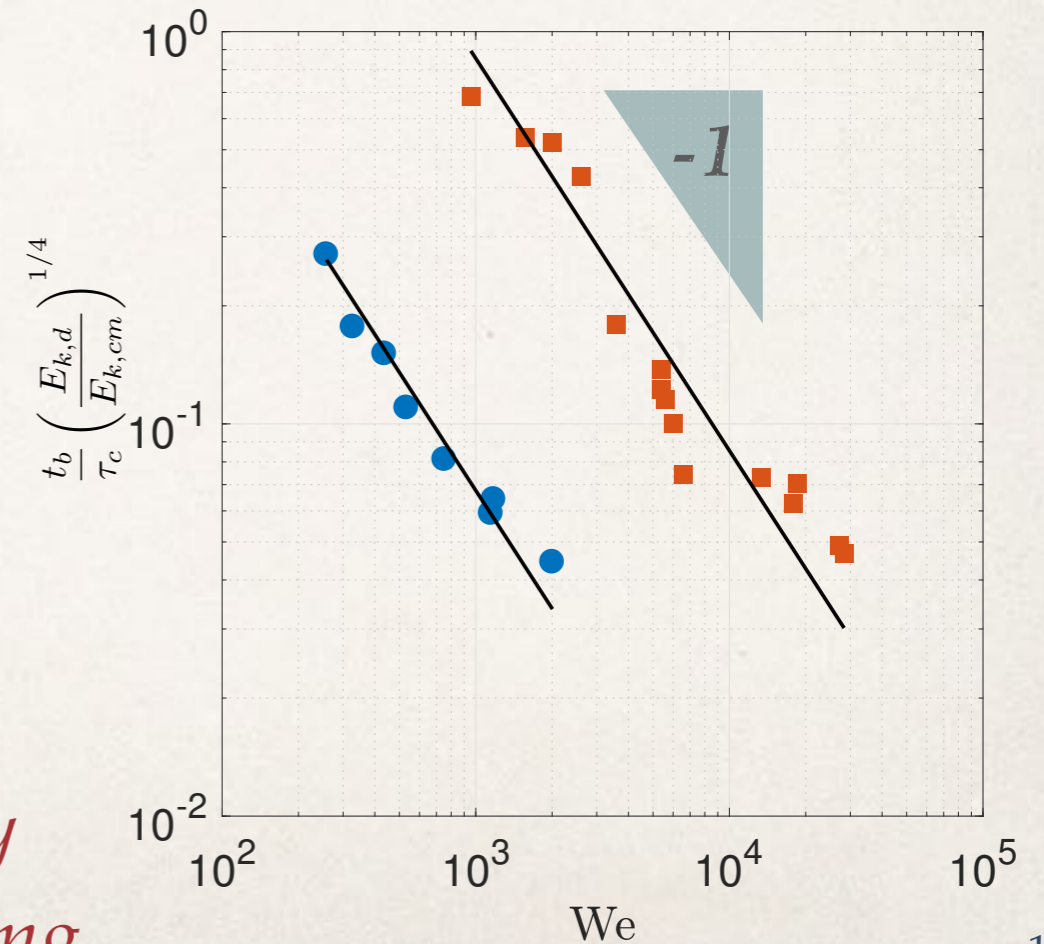
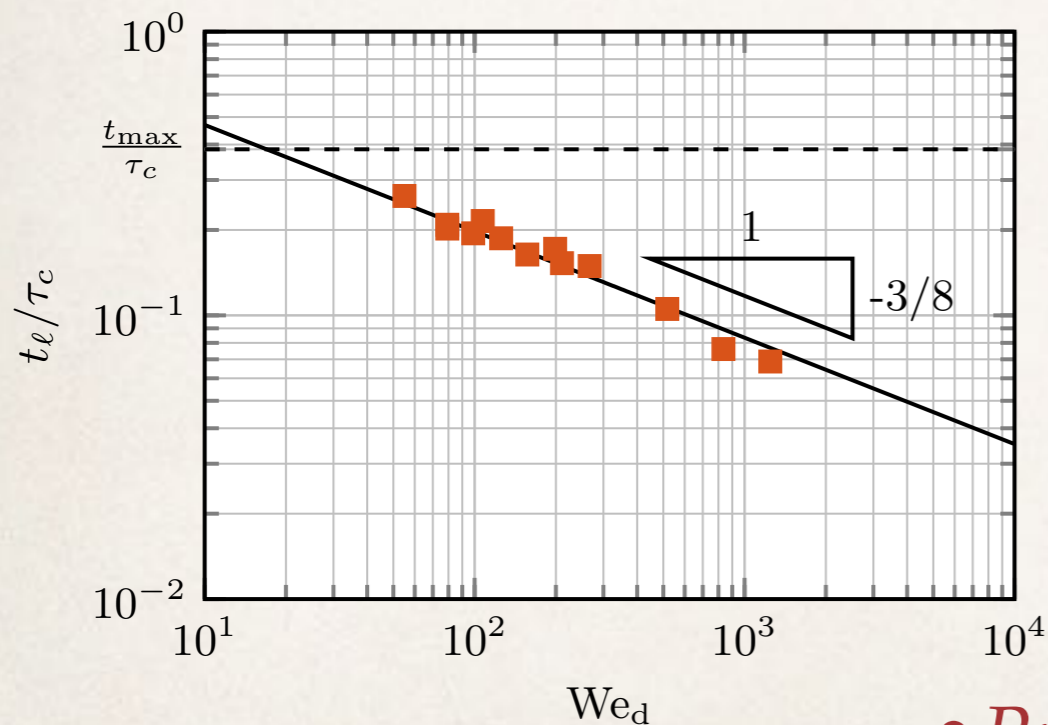
rim breakup

$$t_{\ell}/\tau_c \sim We^{-3/8}$$



hole nucleation

$$t_b/\tau_c \sim We^{-1}$$

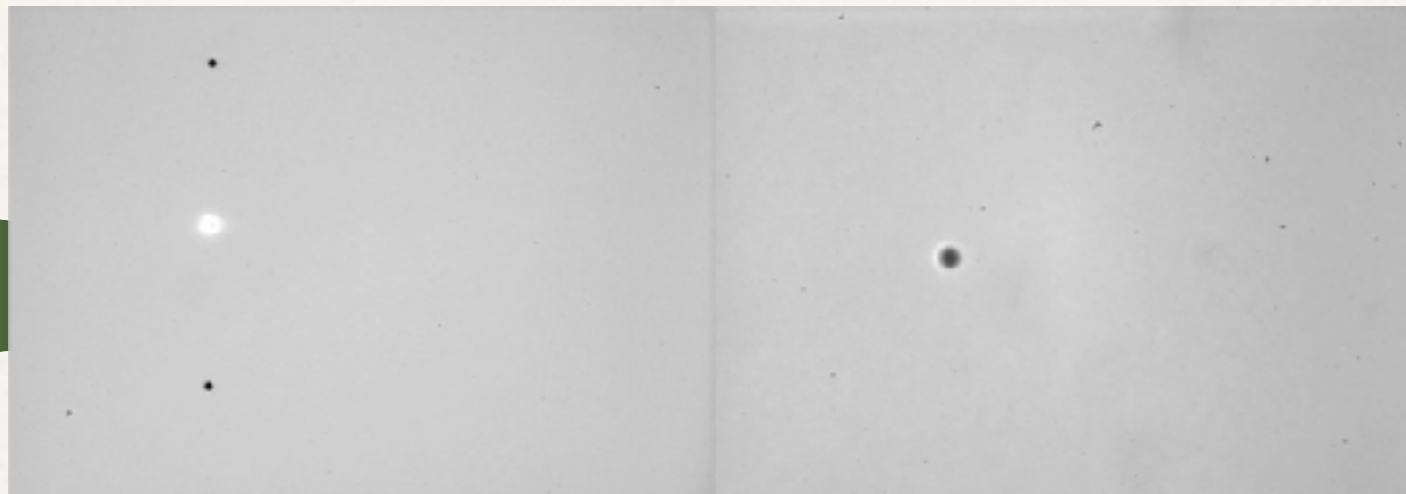


- *Pre-pulse energy*
- *Main pulse timing*

# Drop impact in the EUV source

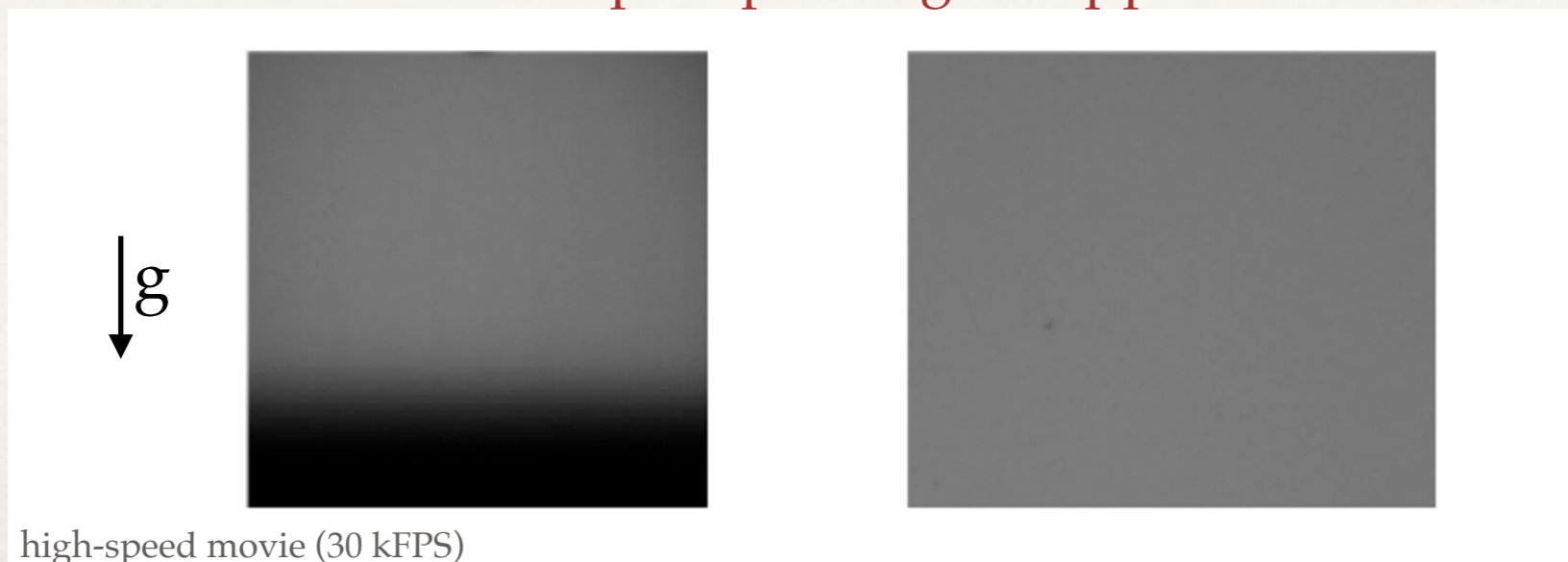
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stroboscopic movie D. Kurilovich & O. Versolato (ARCNL)

2. A 1.6-mm tin drop impacting a sapphire substrate



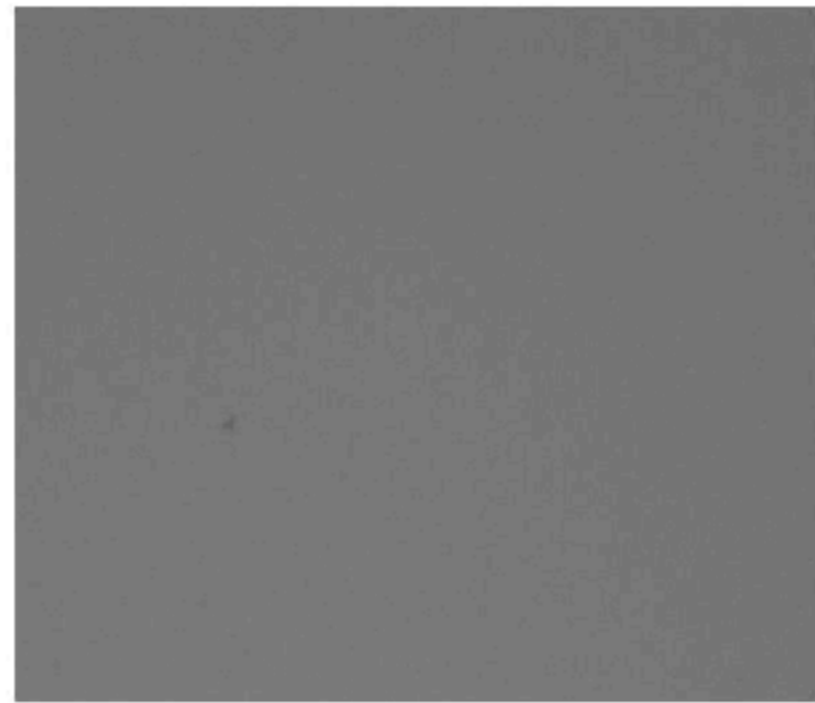
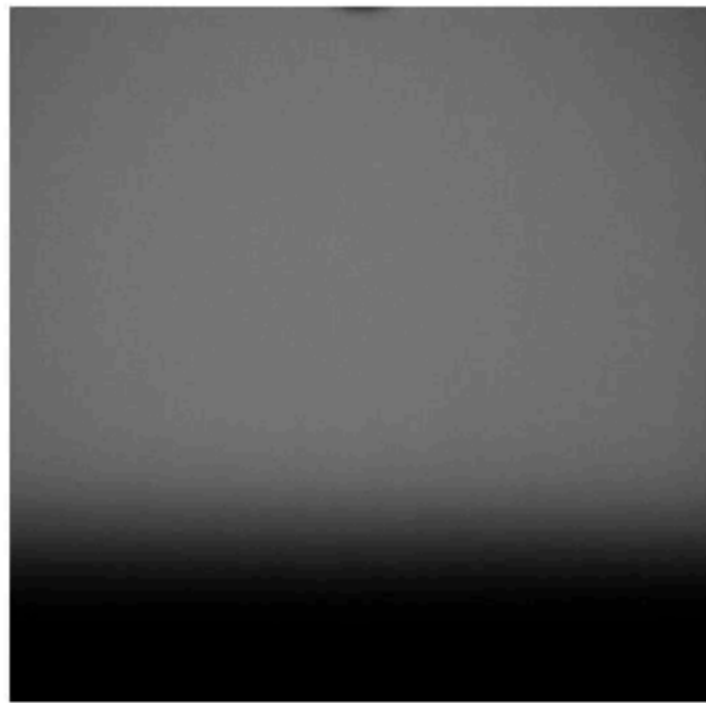
high-speed movie (30 kFPS)

# Isothermal impact of tin on sapphire

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$We = 310$ ,  $U = 3.9$  m/s  
 $T_s = 250$  °C

How does freezing affect impact dynamics?

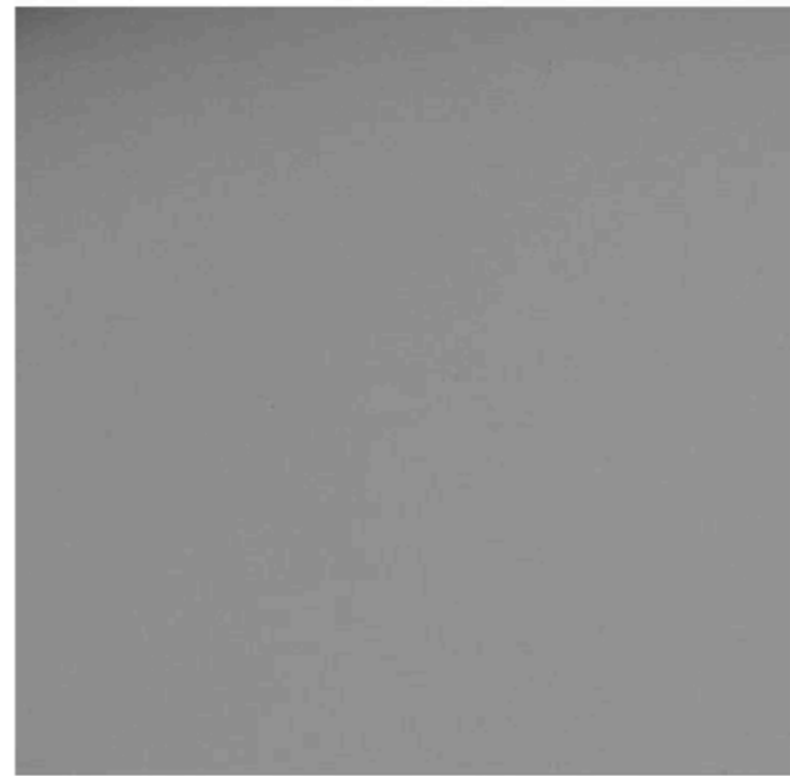
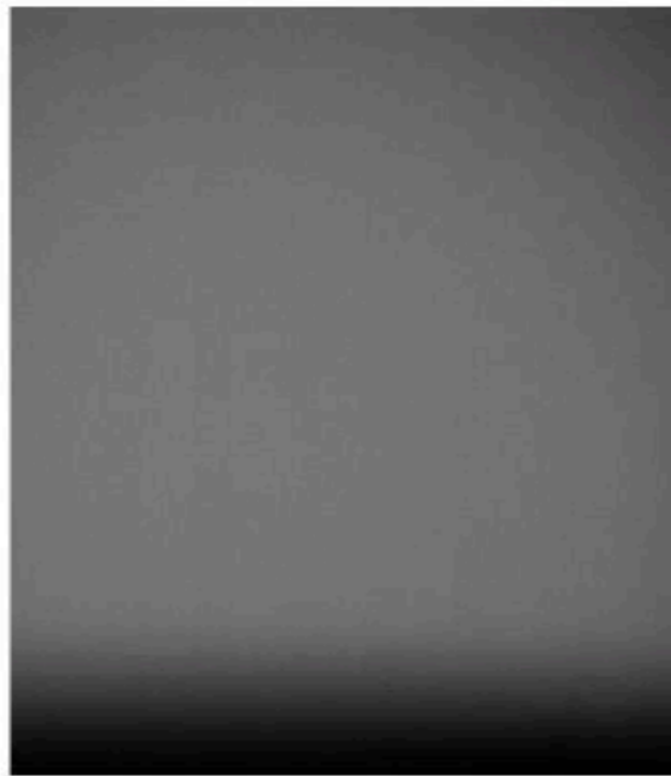


# Tin drop impact on cold sapphire

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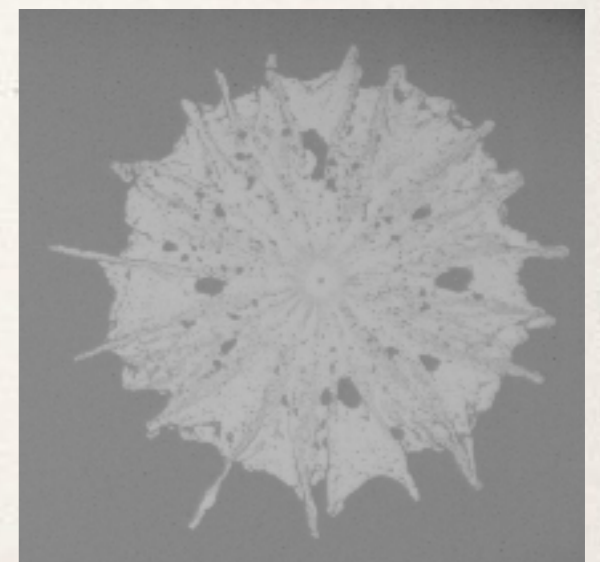
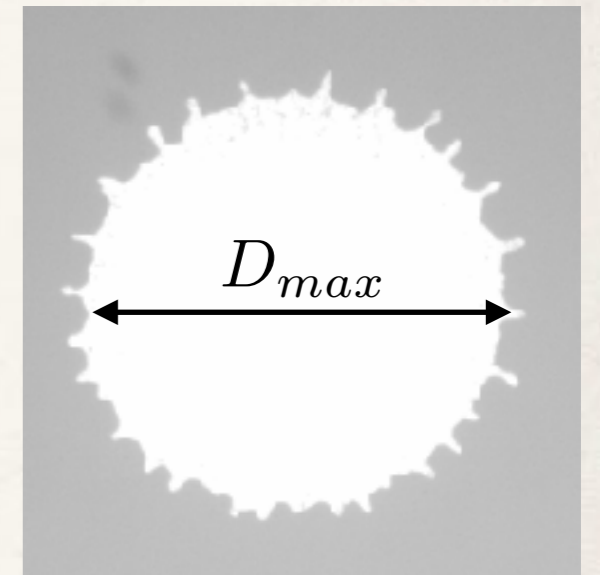
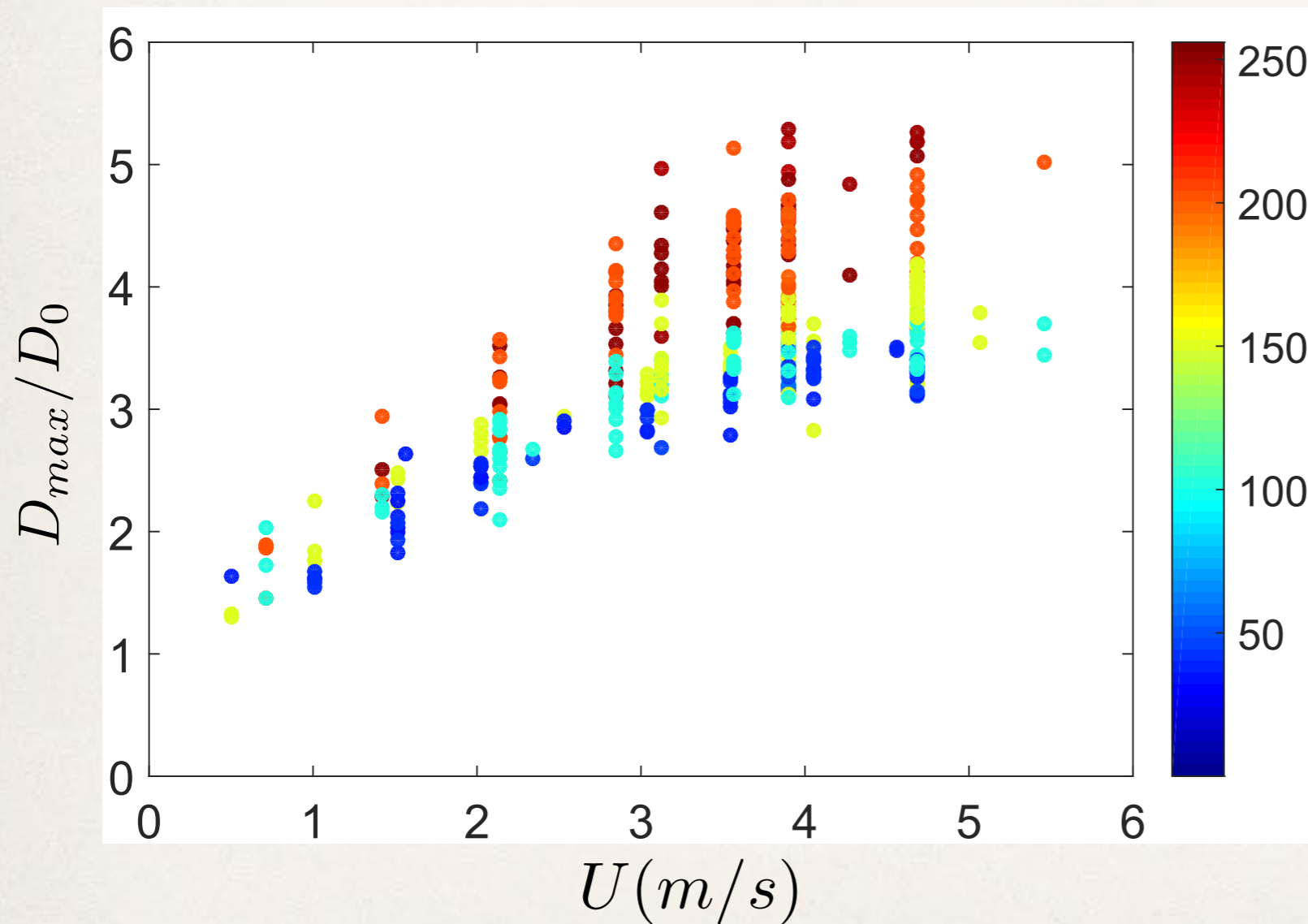
$We = 267$ ,  $U = 3.5 \text{ m/s}$

$T_s = 150 \text{ }^\circ\text{C}$

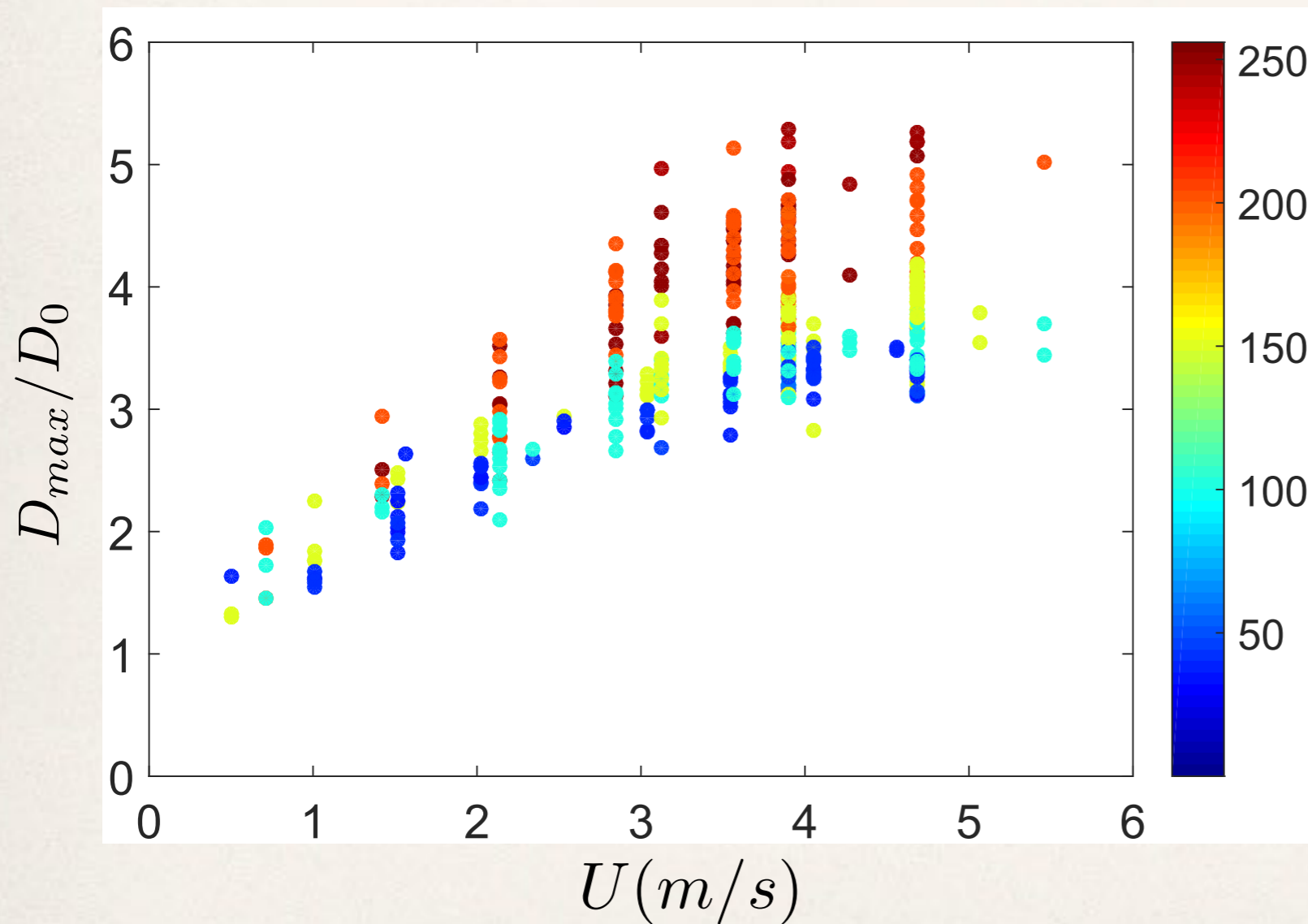


No bounce, no splash, spreading & fingering reduced

# Maximum spreading



# Maximum spreading



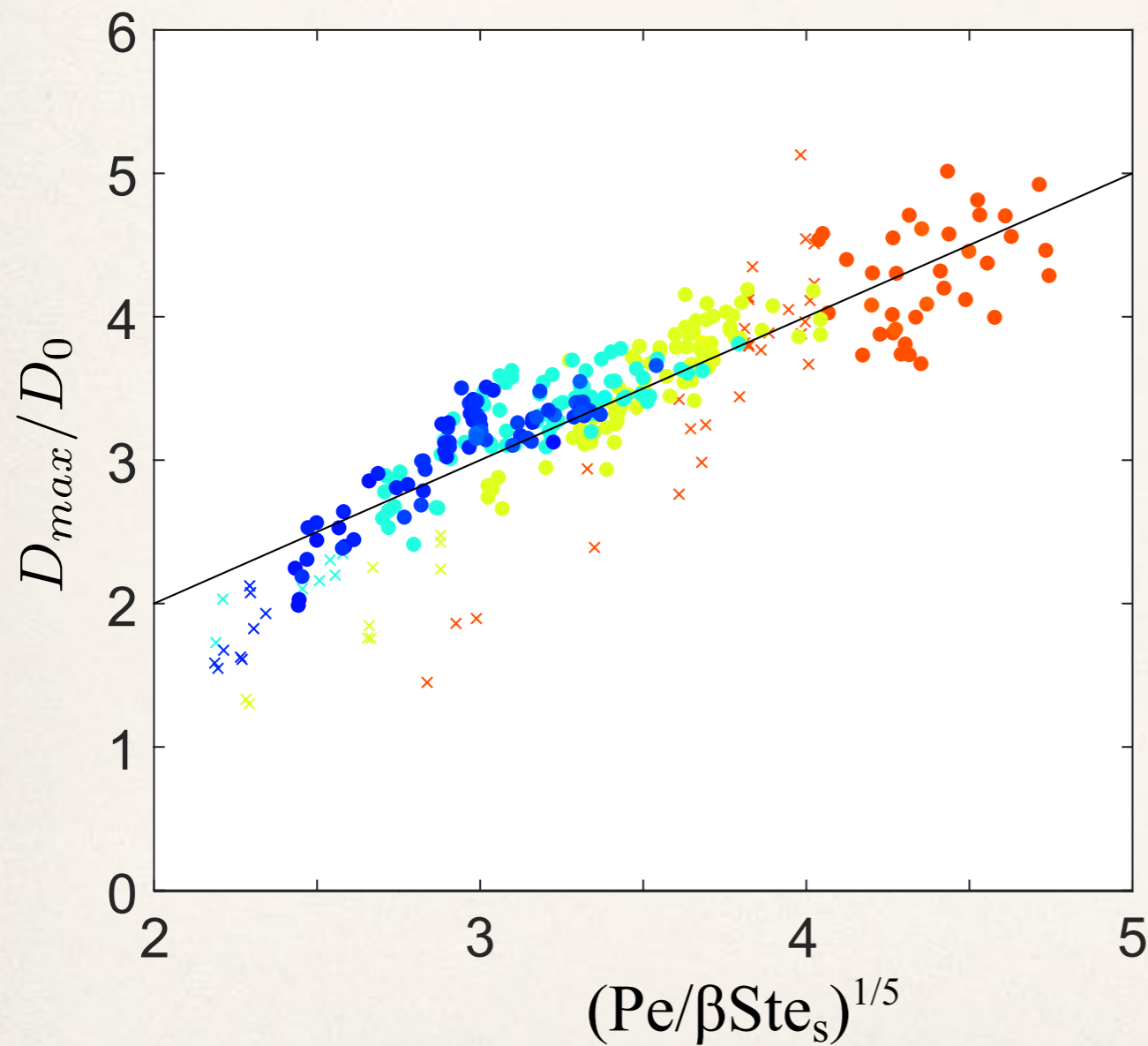
$$\frac{D_{max}}{D_0} \sim \left( \frac{Pe}{\beta Ste_s} \right)^{1/5}$$

$$Pe = \frac{D_0 U}{\alpha}$$

$$Ste = \frac{c_p (T_m - T_s)}{L}$$



# Maximum spreading



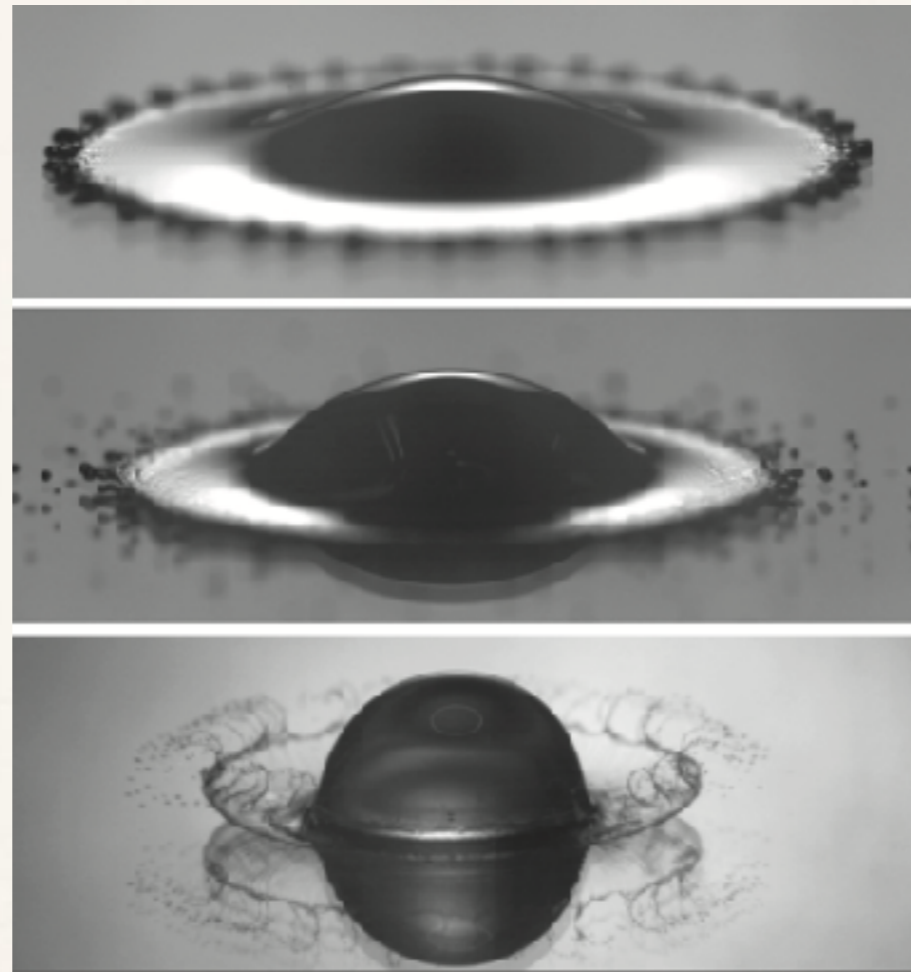
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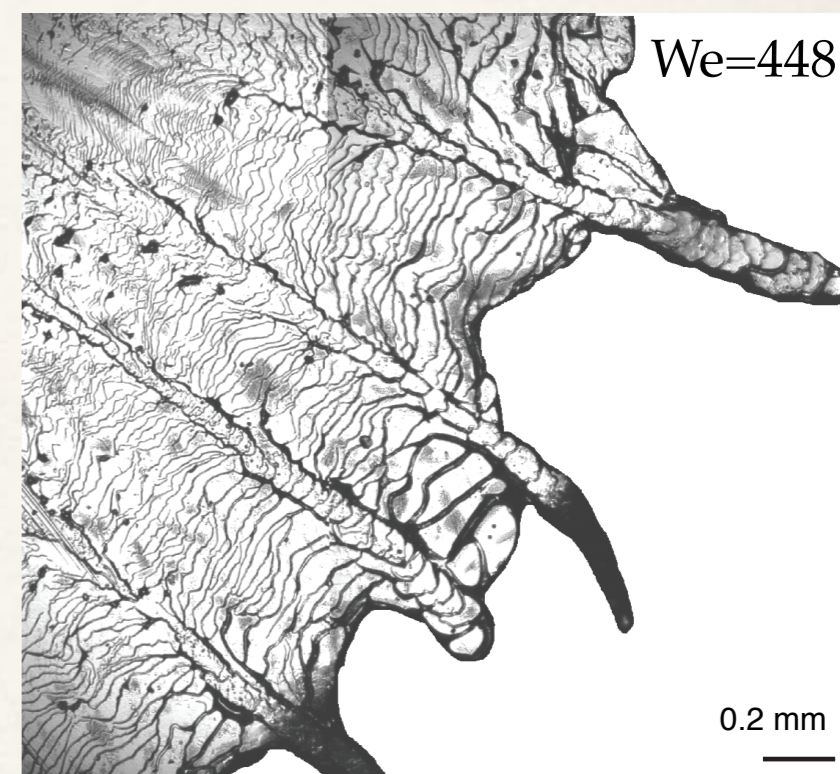
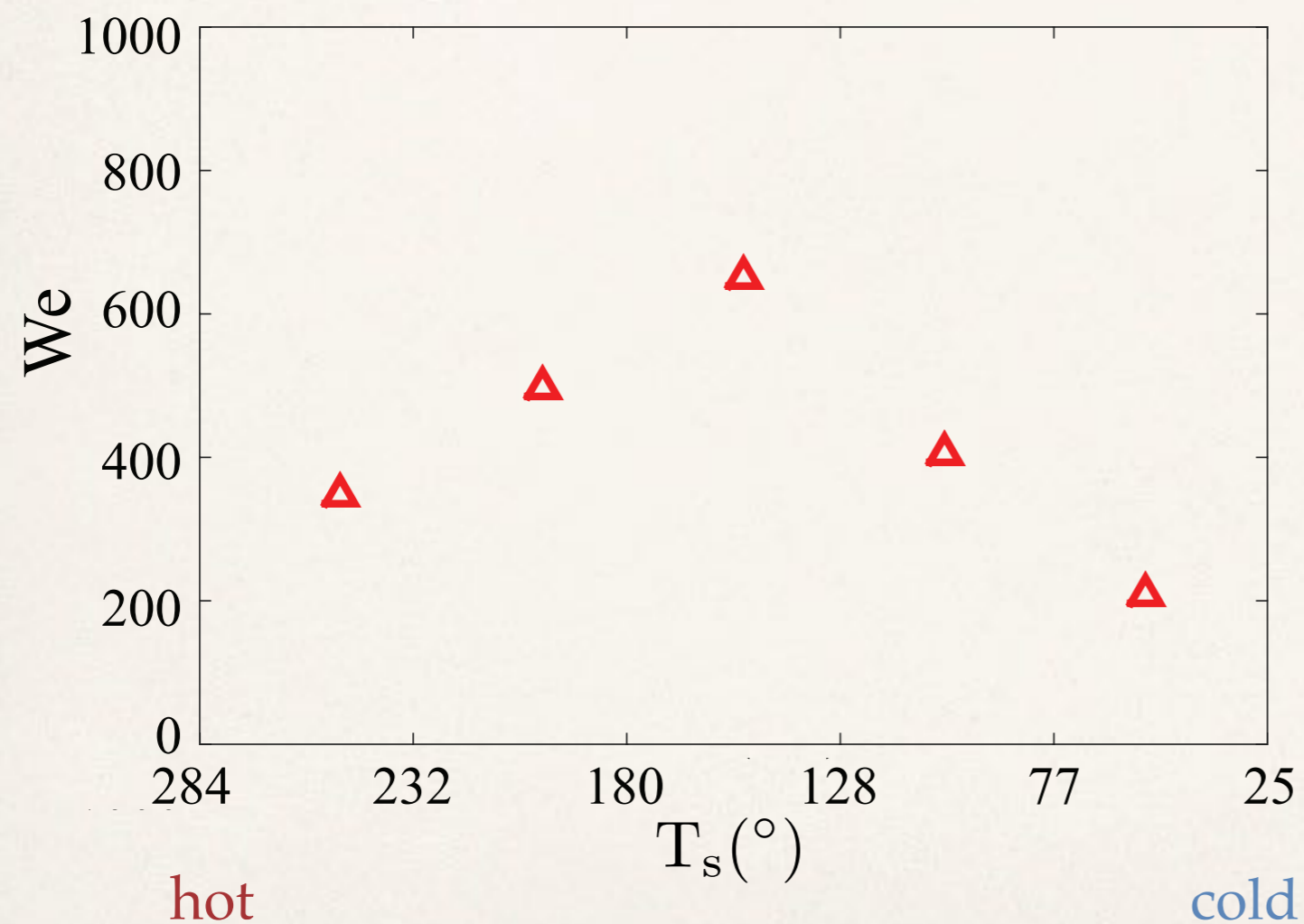
# Splashing threshold

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Josserand & Thoroddsen *Ann. Rev. Fluid Mech.* 48 (2016)

# Splashing threshold

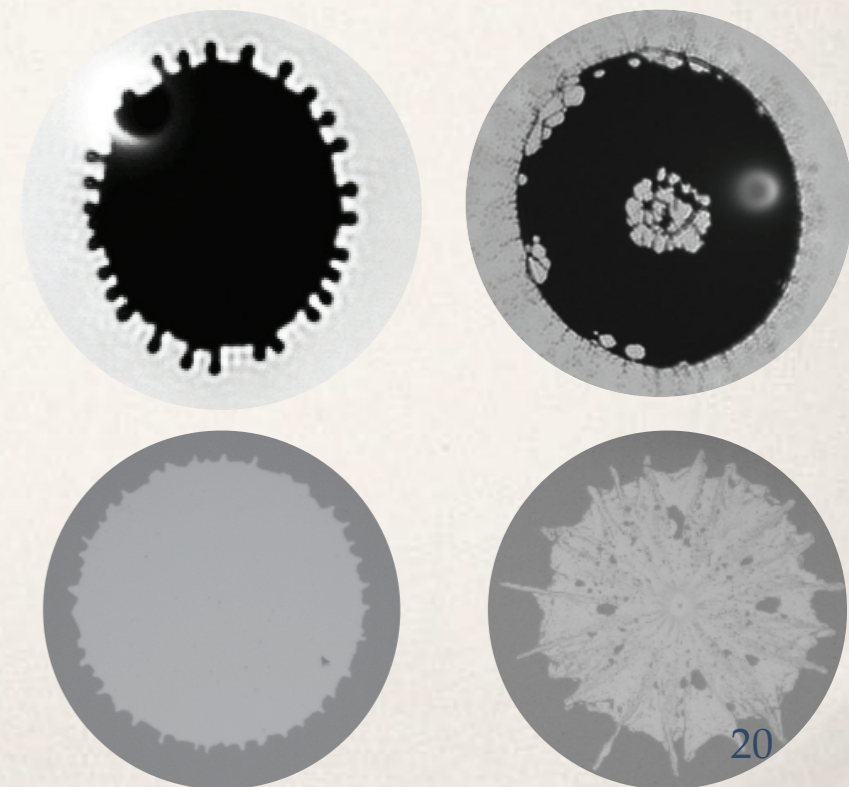


- *Vessel wall temperature*

# Conclusion

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- Impacts by laser pulse & onto a substrate show similar features
- Fluid dynamic instabilities cause drop breakup
- Solidification (cold substrate)
  - ❖ Reduces spreading
  - ❖ Reduces or enhances splashing



# Thank you for your attention!



## Acknowledgements

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*ARCNL*

**Dmitry Kurilovich**

**Bo Liu**

**Francesco Torretti**

**Oscar Versolato**

## More information\*:

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